Automated seismic waveform location using Multichannel Coherency Migration (MCM)–II. Application to induced and volcano-tectonic seismicity

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6 SUMMARY

Locating microseismic events is essential for many areas of seismology including volcano and earthquake monitoring and reservoir engineering. Due to the large number of microseismic events in these settings, an automated seismic location method is required to perform real time 9 seismic monitoring. The measurement environment requires a precise and noise-resistant event 10 location method for seismic monitoring. In this paper, we apply Multichannel Coherency Mi-11 gration (MCM) to automatically locate microseismic events of induced and volcano-tectonic 12 seismicity using sparse and irregular monitoring arrays. Compared to other migration-based 13 methods, in spite of the often sparse and irregular distribution of the monitoring arrays, the 14 MCM can show better location performance and obtain more consistent location results with 15 the catalogue obtained by manual picking. Our MCM method successfully locates many trig-16 gered volcano-tectonic events with local magnitude smaller that 0, which demonstrates its 17 applicability on locating very small earthquakes. Our synthetic event location example at a 18 carbon capture and storage site shows that continuous and coherent drilling noise in industrial 19 settings will pose great challenges for source imaging. However, automatic quality control 20 techniques including filtering in the frequency domain and weighting are used to automati-21 cally select high quality data, and can thus effectively reduce the effects of continuous drilling 22

noise and improve source imaging quality. The location performance of the MCM method for
 synthetic and real microseismic datasets demonstrates that the MCM method can perform as a
 reliable and automatic seismic waveform analysis tool to locate microseismic events.

Key words: Earthquake source observations – computational seismology – time-series anal-

²⁷ ysis – earthquake monitoring and test-ban treaty verification.

28 1 INTRODUCTION

Microseismic or passive seismic monitoring has been used extensively in monitoring geo-industrial 29 applications (e.g., hydraulic fracturing, carbon dioxide storage and mining setting (Power et al. 30 1976; Verdon et al. 2011; Gibowicz & Kijko 2013; Shi et al. 2018a)) as well as hazard monitor-31 ing (e.g., volcano-seismology and slope stability (Wilks et al. 2017; Xu et al. 2011)). As a cost-32 effective monitoring technique, microseismic monitoring is used to demonstrate storage security 33 of carbon capture and storage (CCS) (Verdon et al. 2010; Shi et al. 2018c). It is also an effective 34 method for monitoring volcanoes and forecasting potential eruptions (McNutt 1996; Lavallée et al. 35 2008). Microseismic monitoring can provide geomechanical deformation information induced by 36 fluid injection or flow, which can be used to evaluate rock failure processes in the reservoir of a 37 carbon storage site or volcanic edifice. 38

Noise is an inevitable feature of recorded seismic data. Typically, random noise is assumed 39 to be stationary with a Gaussian distribution, whereas real noise is often non-stationary and so 40 does not conform to a single Gaussian distribution (Birnie et al. 2016; Yuan et al. 2018b). With 41 these features, seismic data with real noise are often more challenging for seismic processing and 42 more difficult to deal with than Gaussian or white noise. For CCS, microseismic monitoring is 43 often conducted during carbon dioxide injection. Therefore, the ambient noise due to the fluid 44 flow and injection exists all the time during the injection process, especially for monitoring arrays 45 which are deployed close to the injection well. Local drilling with associated continuous drilling 46 noise can also affect the recorded seismic data significantly. The injection and drilling noise are 47

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⁴⁸ continuous and are often coherent across many of the receivers. They can form a great challenge
⁴⁹ for microseismic event location in CCS (Barkved et al. 2002; Knudsen et al. 2006; Birnie et al.
⁵⁰ 2016, 2017). Therefore, suitable ways to reduce or remove real noise and obtaining accurate event
⁵¹ location results are required.

For CCS and volcano seismicity, a large number of seismic events can happen within a short 52 period, which can be very difficult and time-consuming to locate by manual arrival time picking 53 (Yuan et al. 2018a). In addition, the ever increasing monitoring data volume and larger monitoring 54 arrays also put great demands on automatic seismic location algorithms for efficient microseismic 55 monitoring. The traditional arrival time based location methods require phase identification and 56 picking, thus are not suitable for automatic event location. Although there are ways to perform an 57 automatic arrival time picking (Bai & Kennett 2000; Maggi et al. 2009), manual picking is still 58 required to increase the picking reliability when the signal-to-noise ratio of seismic data is low or 59 the arrivals of seismic events are overlapped. There have been various migration-based location 60 methods developed to automatically locate seismic events using recorded waveforms (Kao & Shan 61 2007; Gharti et al. 2010; Drew et al. 2013; Grigoli et al. 2013a,b; Zhebel & Eisner 2014; Langet 62 et al. 2014; Cesca & Grigoli 2015; Grigoli et al. 2016). Compared with arrival time based methods 63 where the arrival times are determined by manual picking, automated waveform based location 64 methods do not need phase picking and association, thus are more efficient and have the ability to 65 identify more seismic events. Small, more numerous seismic events which cannot be picked man-66 ually or automatically can be effectively identified by fully utilizing the recorded full waveforms. 67 Thus the automated waveform based location methods can help add more insights into the frac-68 turing process and natural earthquakes. By using the waveforms and the matched filter technique, 69 Peng & Zhao (2009) detected a large number of missing aftershocks along the Parkfield section of 70 the San Andreas fault and used the newly detected seismic events to understand the postseismic 71 deformation around the rupture zone associated with the mainshock of the 2004 Parkfield earth-72 quake. However, the matched filter technique requires reliable waveform templates. Therefore, 73 this technique is not suitable for research areas where there is no available event catalogue. 74

⁷⁵ Migration-based methods have the potential to be applied as real time location schemes, yet

the location reliability and accuracy of these methods is often unsatisfactory in presence of strong 76 noise. Location accuracy is very important in terms of correctly imaging the fracture process and 77 geometry, which can be used to reveal the source mechanism and deformation orientation. Large 78 location errors during microseismic monitoring of CCS and volcano seismicity may contribute to 79 huge economic loss or larger risk as the injection may be terminated prematurely if the induced 80 fracture length has been exaggerated or volcano activity is underestimated because of mislocation 81 of volcano seismicity. The other problem which often challenges migration-based location meth-82 ods is the station coverage and distribution. Sparse monitoring stations hinder the utilization of 83 waveform coherency for migration-based location methods, which causes poor noise-resistance and location performance. Irregular station distribution will reduce imaging resolution and lead to 85 blurred location results. However, due to the restrictions of the actual deployment environment and 86 cost, practical monitoring arrays are often sparse and irregularly distributed especially for natural 87 earthquake monitoring arrays. Therefore, an automatic and precise seismic location method which 88 can work on sparse and irregular monitoring arrays as well as efficiently with dense and/or regular 89 networks is in great demand. 90

Shi et al. (2018b) proposed a fully automated seismic location method based on waveform 91 coherency. This automated location method utilizes Multichannel Coherency Migration (MCM) 92 and is suitable for locating induced seismicity and natural earthquakes. Different to traditional 93 migration-based location methods which locate the source by stacking waveforms of characteristic 94 functions, MCM calculates the multichannel coherency among stations and stacks the coherency to 95 reveal the source location and origin time. By utilizing multichannel waveform coherency, MCM 96 exhibited excellent location performance with high resolution and outstanding noise resistance. 97 The multichannel coherency has also been utilized to improve the horizontal imaging resolution 98 in seismic interpretation (Yuan et al. 2017). Compared to traditional migration-based location 99 methods, MCM can extract more effective information from seismic waveforms, which give it the 100 ability to locate microseismic and resist interference with noise and other non-related events. The 101 theory and synthetic tests of the multidimensional MCM event-location method can be found in 102 Shi et al. (2018b). Here, we demonstrate that the MCM location method can be used to automati-103

cally locate both injection induced and volcano-tectonic microseismic events especially when the 104 monitoring array is sparse and/or irregularly distributed. We also compare and discuss the loca-105 tion results with other commonly used migration methods under different real noise levels using 106 sparse and irregular monitoring arrays. First, as a feasibility study, we use the MCM to locate two 107 volcano-tectonic earthquakes at the Uturuncu Volcano in Bolivia using a sparse monitoring array 108 and also compare the location results with published event locations in the catalogue. We then 109 apply the MCM to automatically locate triggered earthquakes following the M_w 8.8 Maule earth-110 quake at Uturuncu (Jay et al. 2012) using four hours of continuous waveform data. Then, synthetic 111 seismic data of an irregularly distributed monitoring array with real drilling noise were used to 112 evaluate the location performance of different methods for induced seismicity. In order to obtain a 113 satisfactory location result, quality control methods to remove the coherent drilling noise are ex-114 plored and discussed. Finally, location performance and imaging resolution in different directions 115 of different migration-based methods are analysed and discussed in detail. 116

117 2 THEORY AND COMPUTATIONAL EFFICIENCY ANALYSIS

In this section, we will briefly introduce the 2-dimensional MCM (for a more detailed description and the multidimensional MCM see Shi et al. (2018b)). For MCM, at a particular imaging point k and origin time t_0 , the correlation coefficient between the waveforms of two different stations is calculated by:

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$$r_{ij} = \frac{\sum_{t=t_0}^{t_0+t_w} \left[d_i(t+t_{ki}) - \overline{d_i(t+t_{ki})} \right] \left[d_j(t+t_{kj}) - \overline{d_j(t+t_{kj})} \right]}{(N_t - 1) \sigma_i \sigma_j},$$
(1)

where r_{ij} is the correlation coefficient (i.e. coherency) between the waveforms at station *i* and *j*, d_i and d_j are the two input waveforms within the selected time window for station *i* and *j*, t_w is the coherency analysis time window for a particular seismic phase, N_t is the number of time samples in the time window, t_{ki} and t_{kj} are traveltimes of a particular seismic phase from imaging point *k* to the station *i* and *j*, σ is the standard deviation of the corresponding signal and the overlines denote averages.

After calculating correlation coefficients for all possible station pairs, the stacking function can

130 be expressed as

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$$p(x, y, z, t_0) = \frac{1}{N(N-1)} \left(\sum_{i < j}^{N} |r_{ij}^P| + \sum_{i < j}^{N} |r_{ij}^S| \right).$$
⁽²⁾

where r_{ij}^P and r_{ij}^S represent the waveform coherency of P- and S-phases for station pair ij, N is the number of stations and the number of unique receiver pairs equals N(N-1)/2, $p(x, y, z, t_0)$ is the final 4D imaging function and stores the stacked waveform coherency at position (x, y, z) and origin time t_0 (Shi et al. 2018b).

The 4D migration volume contains all the information about source location and origin time. Locations (x_s, y_s, z_s) and origin times t_{0s} of seismic events can be identified by finding the maximum value above a preset coherency threshold within certain time periods

$$p(x_s, y_s, z_s, t_{0s}) = max_{t_0 \in [t_1, t_2]} \{ p(x, y, z, t_0) \ge p_c \}.$$
(3)

As an automated seismic location method, only a few input parameters, i.e. length of coherency 140 analysis time window t_w and coherency threshold p_c , are required for MCM in an event location 141 process. The length of coherency analysis time window t_w should be equal to or larger than the ap-142 proximate period of seismic phases (Shi et al. 2018b). A longer time window is suggested in order 143 to suppress the interference of noise and other incoherent phases when seismic data contain strong 144 noise or coda waves. The coherency threshold p_c is determined according to the background noise 145 level. A higher coherency threshold can help identify seismic events which have more probability 146 to be real seismic events, but will also decrease the number of identified seismic events. It is worth 147 noting that the migration process and the event identification process are two totally independent 148 processes. So it is easy to adaptively adjust the coherency threshold according to a migration vol-149 ume and choose a suitable threshold which can fulfil the requirements of the application. 150

For event locations based on manual picking, the computational efforts depend on the number of earthquake events in the time period. The more events there are, the more expensive it is to pick and locate them. However, for MCM, the computational cost is independent of the number of events. The computational cost is only related to the number of imaging points, the number of searched origin times and the number of stations. The whole MCM procedure is highly parallelizable, and the migration process is quite independent on different scales (from imaging point

level to origin time level). Therefore, parallel computing in MCM can be performed on different 157 imaging points or different origin times according to actual requirements. Very little communica-158 tion is required for MCM when performing parallel computations, e.g. maximum migration values 159 of different origin times when performing parallel computing on different origin times or migra-160 tion values of different imaging points when performing parallel computing on different imaging 161 points. We implement MCM using the Message Passing Interface (MPI) and analyse its compu-162 tational efficiency on a high performance cluster (Figure 1). Both P- and S-waves are used in the 163 MCM calculation and the number of time samples within the P/S time window is 100. Figure 1 (a) 164 shows the computational times for different numbers of imaging points (Ns) and origin times (Nt) 165 used in the MCM. As can be seen in the figure, the computational cost increases linearly with the 166 number of imaging points and origin times, which demonstrates that the MCM workload scales 167 essentially perfectly. Figure 1 (b) shows the computational times for different numbers of stations 168 (N) used in the MCM. As we can see in the figure, the computational cost increases rapidly with 169 the number of stations. Actually, the computational cost is proportional to the number of unique 170 station groups: $N \times (N-1)/2$ (as Figure 1 (b) blue line shows), which is in accordance with the 171 theory of MCM (Shi et al. 2018b). Figure 1 (c) shows the computational times and speedup ratios 172 when different numbers of computing cores (Nc) are used. As expected the computational times 173 (black line) decrease dramatically when more cores are used in the computation. The speedup 174 ratios (blue line) are calculated by dividing the computational times of different cores by the com-175 putational time of a single core. Due to the high scalability of the MCM process, the speedup ratio 176 of MCM is very close to the theoretical speedup ratio (red dashed line). Accordingly, we assume 177 that the computational time is proportional to the number of imaging points, the number of origin 178 times and the number of unique station pairs, and the speedup ratio equals the theoretical speedup 179 ratio. Therefore, the computational time $t = k \times Ns \times Nt \times N \times (N-1)/Nc$, where k is a co-180 efficient related to computer architecture. Using the data of Figure 1 (a-c), we obtain a coefficient 181 of $k = 1.5 \times 10^{-7}$ second with the current settings. If real time processing is required, the MCM 182 calculation time should be less than the length of the data. Here, we assume that the sampling in-183 terval for searched origin times is 0.1 second, and thus we have 10 origin times to process for each 184

second of recorded seismic data. Therefore, for a real time processing, the required cores should 185 fulfil $Nc \ge 10 \times k \times Ns \times N \times (N-1)$. Figure 1 (d) shows the required cores for real time 186 processing when different numbers of stations and imaging points are used in MCM. The real time 187 processing is expensive, but is still feasible with the current computer resources when the number 188 of stations is not very large (e.g. $Ns \leq 40$). For sparse surface monitoring arrays, the number of 189 deployed stations is usually smaller than 20. The number of imaging points can be reduced to less 190 than 300K when locating seismic events in a small region. Therefore, real time processing is com-191 pletely feasible in this situation. For example, for the Uturuncu dataset which we will discuss in 192 detail in the next section, 68 cores are needed to conduct real time processing (shown as the red dot 193 in Figure 1 (d)). Here, we only implement MCM using CPUs and MPI. Because the whole MCM 194 process is highly parallelizable and the computation of MCM can be simultaneously processed in 195 large blocks of data, we anticipate much larger speedup ratios when using Graphics Processing 196 Units (GPU), which we are currently exploring. 197

3 LOCATION OF SHALLOW SEISMICITY AT UTURUNCU VOLCANO

Uturuncu is a long-dormant stratovolcano in Bolivia, which has an elevation of about 6000 m (Jay 199 et al. 2012). Recent studies of surface deformation, fumarolic activity and the earthquake rate of 200 Uturuncu show signs of unrest and potential of eruption again, which calls for close monitoring 201 (Pritchard & Simons 2004; Sparks et al. 2008; Jay et al. 2012). As shown in Figure 2, 15 three-202 component seismometers have been temporarily deployed surrounding the inflating Uturuncu from 203 April 2009 to April 2010 (Pritchard 2009). The farthest station is located about 25 km from the 204 volcano summit. The seismometers have a sampling rate of 50 samples/s, which means the highest 205 effective frequency of the recorded data is 25 Hz. Nine seismometers are short-period instruments 206 and six seismometers are intermediate-period instruments. The tectonic setting of Uturuncu and 207 the catalogue for these events located by manual picking can be found in Jay et al. (2012). We 208 apply the MCM on the recorded continuous waveform data to show the potential of this method in 209 a volcano-tectonic settings, using a sparse seismic network common in such environments. 210

3.1 Locating two local volcano-tectonic microearthquakes

First, we apply four different waveform migration methods to locate two local volcano-tectonic 212 earthquakes at the Uturuncu and compare the location results. The magnitudes of these two local 213 volcano-tectonic earthquakes are below M_L 1.0. The depths of the two shallow volcano-tectonic 214 earthquakes are above the sea level. We use four different waveform migration techniques, i.e. 215 envelope (Kao & Shan 2007; Gharti et al. 2010), STA/LTA (Drew et al. 2013; Grigoli et al. 2013b), 216 kurtosis (Langet et al. 2014) and MCM (Shi et al. 2018b), to compare the performance in this 217 setting. For STA/LTA migration, the short-term time window has been chosen to be 4 seconds 218 and the long-term time window is 40 seconds. The time window for calculating kurtosis is 4 s. 219 For MCM, a coherent analysis time window of 6 s and a two-channel based coherency scheme 220 are used to locate the seismic events. The coherency threshold of MCM is set to 0.13. Because 221 the monitoring array is very sparse, we set the weighting factors of all stations to 1, which means 222 each trace is equally treated and used for migration. The spatial and temporal intervals used in 223 the source imaging are 100 m and 0.08 s respectively. Because the vertical component data show 224 distinct arrivals of P-waves, we only utilize the direct P-wave to conduct MCM for the vertical 225 component data. Similarly for the north-south and east-west components, we only utilize the direct 226 S-wave to image the events. The coherency of the three component data are then added together 227 to obtain the final imaging values of a particular origin time and space point. The location results 228 of the migration methods are compared to the locations in the catalogue. The velocity model used 229 in the event location is the same layered model as described in Jay et al. (2012). 230

Figure 3(a) shows the recorded three-component waveforms at station UTCA for the first event, 231 whose local magnitude M_L is 0.63 (Jay et al. 2012). The direct P-wave and S-wave of this event 232 are distinguishable in the recorded waveforms, but the waveforms contain extended coda. The 233 whole waveform train containing direct waves and coda waves for this event is about 6 seconds. 234 Figure 4 shows the vertical and horizontal profiles of the migration results for the four different 235 waveform migration methods using all available data. The depth (Z-axis) is measured relative to 236 the sea level. The layer at depth 0.5 km which shows a velocity increase can be seen clearly in the 237 migration profiles. The catalogue location (Jay et al. 2012) of this event is displayed as a star in 238

the figure for comparison. For the envelope and STA/LTA migration, the source energy is not well 239 focused. Thus the event location results of these two methods are not reliable, probably because the 240 envelope and STA/LTA cannot identify the event onset from the recorded waveforms. For kurtosis 241 and MCM, the source energy is well focused, thus the location results are more useful. The event 242 location result of the MCM shows better agreement to the location in the catalogue. The location 243 deviations of the MCM result relative to the event in the catalogue are 0.584, 0.557 and 0.469 km 244 in the X, Y and Z directions, respectively (Table 1). Figure 5(a) shows the stacking function of 245 the MCM method at the position of the most coherent point. The stacking function jumps to the 246 maximum value at about one coherent analysis time window earlier than the published origin time 247 of the event and drops down to the noise level quickly. The estimated origin time of the MCM 248 method can be determined from the maximum coherency time, the analysis time window and the 249 period of the direct waves. This is in agreement with Shi et al. (2018b). We will discuss this later 250 in detail in the discussion section. The maximum coherency value is only about 0.16. A longer 251 coherent analysis time window tends to decrease the overall waveform coherency as more data 252 including noise are put into the coherent analysis. However, a longer time window is beneficial 253 for obtaining a stable migration result. The coherency of the coda wave is also included to benefit 254 the source imaging. For this volcano earthquake dataset, tests show that the analysis time window 255 needs to be at least 1 s to eliminate the influence of the noise and pure coda waves. We used a time 256 window of 6 s for both events to make the migration results more stable. 257

Table 1 shows the quantitative location results of the different migration methods and the com-258 parison with the catalogue location. The origin time of this event for the MCM method in the table 259 is estimated using the maximum coherent time plus the coherent analysis time window following 260 Shi et al. (2018b). The event location of the MCM method shows the best correlation to the event 261 location in the catalogue with less deviation in the location and origin time. Predicted P- and S-262 wave arrival times for this event in the catalogue and the event located by MCM are compared on 263 record sections in Figure 6. The direct P- and S-wave arrivals correspond well with the predicted P- and S-wave arrival times for MCM location in most stations. Therefore the location determined 265 by MCM of this event is acceptable. 266

Figure 3(b) shows the recorded three-component waveforms at station UTCA for the second 267 event, whose local magnitude M_L is -0.29. The direct P-wave can be well identified in the vertical 268 component and the direct S-wave can be well identified in the north-south and east-west com-269 ponents. The coda waves following the direct P- and S-waves are obvious. Figure 7 shows the 270 vertical and horizontal profiles of the migration results for the four different waveform migration 271 methods. As with the migration results of the previous event, the envelope and STA/LTA migra-272 tion methods do not focus the source energy appropriately. The migration results of the kurtosis 273 and MCM method are quite similar. The horizontal locations of this event using the kurtosis and 274 MCM method are consistent with the catalogue location with only little deviation. However, the 275 located event depths of both kurtosis and MCM method are deeper than the event depth in the 276 catalogue (1.72 km and 1.92 km deeper respectively). Nevertheless, compared to the horizontal 277 location of the seismic event, the event depth is often not well constrained by the recorded data 278 especially for surface arrays. The trade-off between event depth and origin time often makes event 279 depth determination problematic and more difficult (Eisner et al. 2010). Figure 5(b) shows the 280 stacking function of the MCM method at the position of the most coherent point. Table 2 shows 281 the quantitative location results of the different migration methods and the comparisons with the 282 catalogue. The location results of the MCM correspond very well with the catalogue in the hori-283 zontal directions (with very small deviations of 0.166 km and 0.181 km in the X and Y directions, 284 respectively). Predicted P- and S-wave arrival times for this event in the catalogue and the event 285 located by MCM are further compared on record sections in Figure 8. Probably because of the 286 strong heterogeneity in the subsurface, the recorded waveforms at some stations are not very co-287 herent with the waveforms at other stations. However, the migration results of the MCM method 288 are not seriously affected and seem still reliable. From the record sections of the vertical com-289 ponent (Figure 8 first row), we can clearly see the recorded direct P-wave arrivals show better 290 consistency with the theoretical P-wave arrival times in the record section of the MCM method. 291 This further demonstrates the reliability of the MCM location results. 292

The Uturuncu example shows that MCM can be used as a practical and precise seismic location method for automated volcano-tectonic and natural earthquake monitoring. As MCM utilizes

the waveform coherency across different stations, it performs better under high noise conditions and can obtain a more accurate location result compared to other migration-based location methods. Sparse monitoring arrays will decrease imaging resolution and cause location uncertainties. The utilization of multichannel coherency information in MCM can greatly expand available information used for location and improve imaging resolution (Shi et al. 2018b), which is critical for seismic event location using sparse monitoring arrays.

301 3.2 Locating triggered events on four hours of continuous waveform data

The M_w 8.8 Maule earthquake on 27 February 2010 (at 06:34 UTC) triggered hundreds of earth-302 quakes at Uturuncu with the passage of surface waves and the overtone phases of surface waves 303 (Jay et al. 2012). Those triggered seismic swarms are recorded by the deployed Uturuncu moni-304 toring arrays. According to Jay et al. (2012), the triggered events occurred with the onset of the 305 Love and Rayleigh waves, and the earthquake rate reaches a maximum value of two events per 306 minute with the passage of the Rayleigh wave overtones. We apply the MCM to automatically lo-307 cate these triggered earthquakes using four hours of continuous data (06:00:00 to 10:00:00 UTC), 308 which recorded most of the triggered events. The recorded waveform data at station UTCA are 309 shown in Figure 9. As shown in the enlarged part of Figure 9, there are many small magnitude 310 events which can be very difficult and time consuming to pick manually. As the triggered earth-311 quakes start immediately after the surface wave train, many events occurred in a short time period 312 with very close or overlapping waveform trains. Therefore, it will be very difficult to pick and 313 associate different phases to a particular event. In addition, because of interference of noise and 314 coda waves, it is also very difficult to accurately pick the P- and S-wave arrival times of small 315 seismic events. The manual picking accuracy is highly dependent on human experience. The man-316 ual picking errors will inevitably cause location errors. As the MCM does not require picking and 317 phase identification and the location accuracy of the MCM does not depend on event magnitude 318 (waveform amplitude), it is very suitable to be used to automatically locate those dense triggered 319 microseismic events. 320

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The surface waves and surface wave overtones of the M_w 8.8 Maule earthquake not only

trigger many seismic events in this area but also forms a big challenge for migration imaging 322 using waveforms. Here, we filter waveforms using a frequency band of 4.2 - 21.6 Hz to exclude 323 the influence of surface waves and low frequency noise. Because the sample rate (50 samples/s) 324 is low, we suggest to use a long time window for coherency analysis in the MCM. Using a longer 325 time window can improve the imaging stability and quality in noisy situations. We adopt a four 326 second time window for both P- and S-waves in the MCM to resist the interference of noise 327 and coda waves. Similarly, we only utilize the direct P-wave to conduct MCM for the vertical 328 component data, and only utilize direct S-wave for the horizontal component data. The coherency 329 value of the P-wave for the vertical component data and coherency values of the S-wave for the 330 two horizontal component data are then stacked together to form the final migration value. For 331 conventional waveform migration methods which stack amplitudes or characteristic functions of 332 amplitudes, S-phases are often assigned higher weighting factors because S-phases tend to have 333 higher amplitude. However, the S-phase often interferes with coda and converted waves, thus tends 334 to have lower waveform coherency across different stations. In contrast, the P-phase which arrives 335 first often has higher coherency despite its lower amplitude. Therefore, we assign a weighting 336 factor of 0.6 to the P-phase of the vertical component and factors of 0.2 to the S-phases of each of 337 the two horizontal components (east-west and north-south), noting that the MCM is insensitive to 338 amplitudes. 339

The imaging area is 18 km, 15 km and 8 km in north-south, east-west and vertical directions, 340 respectively (as shown in Figure 2 white rectangle area). The imaging point interval is 200 m in 341 all different directions. Therefore, there are 283,556 imaging points in total. The time interval for 342 searching for origin times is 0.08 s. The total number of searched origin times in the four hours is 343 about 180,000. We assume that two earthquakes will not occur at the same time or within a few 344 origin time samples (0.08 s). Therefore, at each searched origin time, we only save the imaging 345 point which has the maximum coherency value. Figure 10 shows the variation of the maximum 346 coherency value with different origin times in the four hours. When an earthquake occurs, at the 347 correct origin time, the coherency values of each imaging point will all rise due to the arrival of 348 the long waveform trains including direct, converted and coda waves. Therefore, we can observe 349

many local peaks rising from the background noise in Figure 10, which potentially correspond to 350 seismic events. We use a coherency threshold of 0.1. By identifying the maximum value of each 351 local peak, we can find the location and origin time of each seismic event. We identify 560 local 352 peaks in Figure 10, which are viewed as potential seismic events. We then check each potential 353 seismic event using the corresponding record sections of these potential seismic events and verify 354 322 seismic events which have clear phase arrivals. The verified seismic events are shown as red 355 dots in Figure 10. Although there are many events which do not show clear P- and S-phase arrivals, 356 they may still be real seismic events, because the signal-to-noise ratio (SNR) for these events may 357 be small (smaller than 1). Since the MCM has the ability to resist strong noise, it is not surprising 358 that it can successfully identify seismic events below the noise level. The problem is that although 359 identified by MCM, the weak seismic events with SNR below 1 cannot be effectively verified 360 through their record sections at the present. By adopting stricter parameters (such as a higher 361 coherency threshold, higher source prominence and longer origin time gap), we can also reduce 362 the number of unverifiable seismic events and improve the proportion of confirmed seismic events. 363 However, this would inevitably result in losing some small real seismic events which cannot be 364 effectively verified by inspection of the record sections at the present. Therefore, further studies 365 about detecting and verifying seismic events (especially events with low SNR) from migration 366 traces/volumes are still needed. Here, since the verifying process is very quick and easy, in order 367 to identify as many seismic events as possible, we adopt relatively relaxed parameters to identify 368 the local maxima in the coherency time slice (Figure 10). 369

The existing catalogue has 114 seismic events in total in this four hour time period in this area, 370 which are located by manual picking. For those 114 seismic events, 112 events (98.25%) have 371 been successfully located by the MCM. In addition, the MCM has also automatically located 210 372 more seismic events than the existing catalogue, which have been verified on the record sections. 373 By checking the corresponding record sections, we find that the MCM not only automatically 374 locates many more triggered seismic events than the catalogue, but also the origin time estimates 375 of most events are more accurate than the existing catalogue under the current velocity models. 376 This demonstrates that MCM is an efficient and reliable automatic location method. Figure 11 377

shows the locations of the 322 verified seismic events and the 114 current catalogue events. In the 378 figure, we can see that the distribution of automatically located seismic events is consistent with 379 that of the events in the catalogue. There are two main earthquake clusters. One is located in the 380 northern part of the study area and close to the volcano. This earthquake cluster occurred earlier (6 381 am to 8 am), and the events are mainly triggered by the surface waves (Love and Rayleigh waves) 382 of the Maule earthquake. The other earthquake cluster occurred from 8 am to 10 am and is located 383 in the southern part of the study area. The seismic events are mainly triggered by the surface wave 384 overtones of the Maule earthquake. 385

Figure 12 shows a seismic event (referred to as event 1) which is both located by the MCM 386 and the manual picking (catalogue). The MCM location result has a similar horizontal location 387 as the catalogue result, but is deeper than the catalogue event. From the corresponding record 388 sections (Figure 13), we can clearly see that the predicted arrival times of the MCM results have 389 a much better correspondence with the P- and S-phase arrivals, especially for the S-phases of the 390 horizontal components. This demonstrates that the MCM location results are reliable and have a 391 better estimation of the origin times of seismic events. Figure 14 shows the migration profiles and 392 record sections of a newly identified seismic event (referred to as event 2) by MCM, which is not 393 in the existing catalogue. The source energy focuses nicely in the migration volume. The record 394 sections which show clear P- and S-wave arrivals also indicate a real microseismic event occurred. 395 It is worth noting that although event 2 is lower in event magnitude and has smaller amplitudes 396 than event 1, the waveform coherency (0.22) of event 2 is higher than that of event 1 (0.17). For 397 MCM, the waveform coherency not only depends on the amplitude (relating to event magnitude 398 and SNR), but is also influenced by the interference of coda waves, converted waves and arrivals 399 from other events. Thus, it is not surprising that a small seismic event can have higher waveform 400 coherency and focussing of migration energy than a larger seismic event when the small event is 401 less affected by interference of other non-coherent waves. This characteristic makes MCM very 402 suitable for locating microseismic events. 403

Many more seismic events have been identified by MCM in this four hour time period than the published catalogue, which greatly complements the catalogue. We provide our extended cat-

alogue in the supplementary material. Figure 15 shows the number of triggered seismic events 406 within the four hours. With this more complete catalogue, we find that rates of triggered events 407 rise shortly after the passage of surface waves or surface wave overtones. Different to Jay et al. 408 (2012), who conclude that rates of triggered events increased to a peak value of two events per 409 minute with the passage of the X2/X3 Rayleigh wave overtones, we find that earthquake rates 410 reach a peak value of about five events per minute after the passage the G1/R1 surface waves (Fig-411 ure 15). An increase of seismicity after the passage of X2/X3 is noticeable, but only reaches about 412 three events per minute. 413

414 4 AQUISTORE SYNTHETIC DATA WITH REAL NOISE

Synthetic waveform data with added Gaussian noise is often used in testing the performance of 415 location algorithms. However in reality, the real noise field is not white, stationary or Gaussian 416 (Birnie et al. 2016). Several noise studies have shown that seismic noise is often variable in space 417 and time, leading to increased difficulty in source imaging (Birnie et al. 2017). In this section, 418 we apply the MCM location algorithm to the Aquistore noise dataset to examine the location 419 performance in the presence of real seismic noise. The Aquistore noise data have been extracted 420 from a permanent surface array installed at the Aquistore carbon dioxide storage site (Roach et al. 421 2015; Birnie et al. 2016, 2017). The monitoring data used here were recorded by the surface 422 array during the drilling and construction phase of the injection and observation wells prior to 423 CO2 injection. Therefore, significant drilling noise and non-stationary noise were recorded in the 424 dataset. No injection-related or induced seismic events are recorded in this period, which makes 425 the recorded time-series an excellent dataset for investigating the effect of real seismic noise on 426 seismic location. Figures 16 (a) and (b) show the surface array geometry and velocity model of 427 the Aquistore area. The surface array consists of 50 buried geophones (34 in North-south direction 428 and 16 in East-west direction) with a sampling frequency of 500 Hz. 429

We generate waveform seismic data using the propagator matrix technique of Zhu & Rivera (2002) for both a shallow and a deep event (Figures 16 (c-e)). The shallow and deep events are located at a depth of 2.55 km and 3.15 km, respectively. The deep event has been placed in a thin

and relatively low velocity layer. There are also many thin layers above and below the deep event 433 (Figures 16 (b-e)), which may cause difficulty in imaging the deep event. We use the shallow and 434 deep events to examine the influence of complex velocity model on the migration result. For both 435 the shallow and deep event, a 45° dip-slip double-couple source with 40 Hz peak frequency is used 436 to give a specified radiation pattern. The recorded real noise (Birnie et al. 2016) is added to the 437 synthetic data to mimic as closely as possible a 'real' dataset with varying signal-to-noise ratios. 438 The SNR is defined by the ratio of the maximum amplitude between signal and noise. This kind 439 of semi-synthetic dataset enables a quantitative evaluation of the location errors in the presence of 440 different realistic noise scenarios and has been employed to evaluate the monitoring performance 441 of a dedicated seismic monitoring array (López-Comino et al. 2017). The synthetic data and noise 442 data are shown in Figure 17. After adding noise, the arrival of the direct P-wave cannot be easily 443 recognized. Stations 18-24 and 41-43 are deployed near the injection and observation well (as 444 shown in Figure 16(a)), and thus are seriously contaminated by drilling noise (Figure 17c). The 445 non-stationarity and spatial variability of the noise will make event location more difficult. 446

447 4.1 Location results for shallow event

We compare the location results of different migration methods using waveform envelope, STA/LTA 448 and kurtosis as characteristic functions and also the MCM method for different SNRs. The same 449 monitored real noise of different levels have been added into the synthetic dataset to make the 450 semi-synthetic datasets of different SNRs. The SNRs are chosen to be infinite (noise free), 1, 0.5, 451 0.25 and 0.025 respectively. We then analyse the influence of SNRs to location results and com-452 pare the performance of different migration methods under different SNR situations. Figure 18 453 compares the migration results for the four different methods when the SNR is 1 (for a complete 454 comparison of different SNRs, see supplementary material Figures S.1-S.3). When the SNR is 455 larger than 0.25, the MCM exhibits the best resolution and location performance in both the hor-456 izontal and vertical directions. Due to the use of the derivative (Langet et al. 2014), the kurtosis 457 seems to have better resolution in the XZ profile (as shown in Figure 18). However the location 458 results of kurtosis migration are often biased due to the trade-off between depth and origin time. 459

The results also show that receiver distribution influences the results of the locations. Compared 460 to the X direction, the image in the Y direction relies on fewer geophones leading to increased 461 location uncertainty in that direction, and therefore the envelope, STA/LTA and kurtosis methods 462 show poorer resolution in the Y direction (as shown in Figure 18). However the MCM still main-463 tains very good resolution in the Y direction. The location results of the envelope, STA/LTA and 464 kurtosis methods are often biased in the Z direction. When the SNR is below 0.25, the MCM fails 465 to locate the source, because the noise recorded during drilling and construction of the injection 466 well is pervasive over all the traces, especially notable in the traces which are close to the injection 467 well. The drilling and construction noise coming from the injection well is continuous in time, and 468 so leads to continuous coherent noise on all the traces. When the SNR is below 0.25, the drilling 469 noise dominates the wavefield in all the traces. The continuous (both in space and time) and co-470 herent drilling noise contributes to the failure of the MCM method when the SNR is below 0.25. 471 The other methods also fail to locate the source accurately because of strong noise contamination. 472 When the SNR is 0.025, all the methods fail to locate the source. However, approaches have been 473 developed to 'whiten' the noise and hence to reduce the influence of coherent noise (Birnie et al. 474 2017). 475

The automatic weighting scheme can be integrated into the multidimensional MCM flexibly 476 (Shi et al. 2018b), which gives MCM the ability to conduct automatic quality control of the input 477 data. We devise an automatic quality control scheme to deal with the drilling noise and surmount 478 the SNR limit in the presence of continuous drilling noise. The automatic quality control scheme 479 comprises weighting and filtering. The weighting factors (Shi et al. 2018b, equation 3) are deter-480 mined by evaluating the amplitude of each trace. Because the continuous drilling noise will nor-481 mally contaminate a whole trace, here we use an average absolute amplitude ratio to discriminate 482 very noisy traces and apply a weighting coherency calculation scheme to all traces. The absolute 483 amplitude ratio of a trace is defined as the ratio of the average absolute amplitude of the trace to the 484 average absolute amplitude of all traces ($a_i = \overline{|\mathbf{d}_i|}/|\overline{\mathbf{D}}|$, a_i is the absolute amplitude ratio of the i-th 485 trace, d_i is the waveform amplitudes of the i-th trace and D is the amplitudes of all traces). Figure 486 19 shows the absolute amplitude ratios of different stations for the noisy datasets with different 487

SNRs. For traces which are highly contaminated by continuous drilling noise, the energy of this 488 trace will be much larger than the average energy over the whole traces, which will contribute to a 489 high absolute amplitude ratio (as shown in Figure 19). Through inspecting the absolute amplitude 490 ratios of all traces, we can identify high quality traces and thus stabilize the migration result. We 491 set an absolute amplitude ratio limit of 1.5. Above this limit, the weighting factor of this trace will 492 be set to 0, otherwise the weighting factors are 1. Because our waveform coherency is evaluated 493 through correlation coefficient, the absolute value of the amplitude will not affect the coherency 494 calculation. Therefore, weighting values of 0 or 1 rather than sliding values are assigned to exclude 495 or include traces in the coherency calculation. Through weighting, we select high quality data to 496 conduct migration and exclude traces with very high absolute amplitude ratio (which means ex-497 tremely low SNR for that trace). As shown in Figure 19, traces 19-24 and 41-43 which are close to 498 the observation and injection wells are highly contaminated by the drilling noise (consistent with 499 Figures 16 and 17). Therefore, after weighting these traces will be excluded from the dataset used 500 for imaging. Before calculating multichannel waveform coherency or characteristic functions, the 501 selected data are filtered in the frequency domain. Because the drilling and construction of the in-502 jection well are low-frequency processes, we applied a 6th-order highpass Butterworth filter with 503 a cutoff frequency of 50 Hz to the semi-synthetic data to remove the low frequency drilling noise. 504

The migration results with automatic quality control scheme (weighting and filtering) are 505 shown in Figure 20 and the SNR before filtering is 0.025 (for a complete comparison of dif-506 ferent SNRs, see supplementary materials Figures S.4-S.6). Through the automatic quality control 507 scheme, the imaging quality of the four migration methods becomes better and the imaging reso-508 lution also improves especially for low SNR scenarios. The MCM exhibits better location results 509 with higher resolution compared to the other methods for all SNR situations. When the SNR is 510 above 0.025, MCM can locate the source accurately without deviation, while the other three meth-511 ods all have location deviations. When the SNR is 0.025, only the MCM can locate the source 512 correctly with a minimal deviation of 20 m. With such a low SNR, the STA/LTA method focused 513 at the shallow part of the true source position with very low imaging resolution, while the kur-514 tosis method cannot focus correctly. Because of the non-Gaussian property of the real noise and 515

the sensitivity of the characteristic function of the kurtosis method, the kurtosis method is more 516 susceptible to the array geometry. An irregular and/or sparse monitoring array will tend to bias the 517 location results of the kurtosis method. Thus the location results of the kurtosis method are less sta-518 ble compared to the other three methods. Figure 21 shows the location errors of the four methods 519 under different SNRs with/without automatic quality control scheme. The MCM method outper-520 forms the other methods at all noise levels when the automatic quality control scheme is applied 521 (Figure 21(b)). The implemented automatic quality control scheme using filtering and weighting 522 can effectively improve the location accuracy for most tested methods. 523

524 4.2 Location results for deep event

The location results for the deep event with a SNR of 1 are shown in Figure 22. Since the SNR 525 is relatively high, for consistency and better comparison with the migration results of the shallow 526 event (Figure 18), original data without automatic quality control are used for migration. The 527 velocity model above the deep event is more complicated as it contains thin layers and large 528 velocity contrasts. However, compared to the shallow event, the location results of the deep event 529 are not seriously affected by the complexity of the velocity model. Due to the increase of the 530 velocity in the imaging area, the arrival time differences between the adjacent imaging points 531 become smaller, which is detrimental for distinguishing the phase arrivals. Correspondingly the 532 imaging resolution for all the 4 methods decreases compared to imaging results of the shallow 533 event (as can be seen in the comparison of Figure 18 and 22). The imaging results of the envelope 534 and STA/LTA methods still have large deviations in the vertical direction, while the MCM and 535 kurtosis methods locate the deep event accurately. The imaging results of the MCM exhibit high 536 resolution in the horizontal direction. However, the resolution in the vertical direction deteriorates 537 compared to the results of the shallow event. The degradation of the vertical resolution is related 538 to the chosen length of the time window of the coherence analysis as well as the velocity of 539 the imaging area. Although the same time window is applied in the imaging of the shallow and 540 deep events, the higher velocity of the deep event layer contributes to the reduction of the vertical 541

resolution. Using a smaller time window can improve the imaging resolution, but at the expense
 of reducing noise suppression ability.

Figure 23 shows the stacking functions of the four methods at the true source location of the 544 deep event. The four methods all exhibit excellent source prominence at the correct origin time. 545 Time windows for both P- and S-phases are simultaneously used in the migration. The pink area 546 around -0.6 second in Figure 23 highlights the time range where P-phases move into the stacking 547 window of the S-phases when searching for origin time. Meanwhile, the pink area around 0.7 548 second highlights the time range where S-phases move into the stacking window of the P-phases. 549 For the stacking functions of the envelope and STA/LTA methods, a notable peak can be observed 550 at these times. However, the MCM can effectively suppress this kind of disturbance and avoid 551 identifying unrealistic events. 552

From the results of the Aquistore dataset, we can see that MCM can be used as an effective 553 migration method to automatically locate microseismic events induced by fluid injection or hy-554 draulic fracturing. Although drilling or injection noise can pose big challenges for source imaging, 555 different ways can be adopted to acquire reliable seismic location results. Irregularly distributed 556 monitoring arrays will lead to unbalanced imaging resolution in different directions. However, due 557 to the utilization of multichannel waveform coherency, MCM can acquire higher and much more 558 balanced imaging resolution in different directions compared to other migration-based methods. 559 As traveltime differences between adjacent imaging points in low velocity zone are larger, the 560 imaging results in the low velocity zone (i.e. a shallow event) are better than those in a high veloc-561 ity zone (i.e. at greater depths), and the source imaging resolution in the low velocity zone is also 562 higher than that in the high velocity zone. 563

564 5 DISCUSSION

For the synthetic data case, the stacked coherency trace will normally exhibit a flat top as we have discussed in detail in Shi et al. (2018b). However, for real data, such flat tops may not exist because of strong interference from noise and coda waves (as shown in Figure 5). Typically, one records and takes the time and position which has the maximum coherency value as the origin time and

596

location of a seismic event. However due to a systematic bias between the origin time and the 569 maximum coherent time, calibration is needed in order to obtain an accurate estimation of the 570 origin time of the seismic event. As shown in Figure 24, similar to synthetic data, the coherency 571 will start to rise at one coherent analysis time window (referred to as Tw) before the correct source 572 origin time (referred to as T0). For real seismic data, because of subsurface heterogeneity, there 573 are many coda waves following the direct P- and S-phases. Those coda waves often show lower 574 coherency compared to direct phases and have high amplitudes compared to background noise. 575 Here, we assume the coda waves are incoherent. Therefore, the maximum coherency value will 576 appear one period (referred to as T) of the direct phase after the rise of the coherency. That is to say 577 the maximally coherent time (referred to as Tmax) is Tw - T ahead of the correct source origin 578 time. So the calibration equation for source origin time is $T_0 = Tmax + Tw - T$. Since it is easy to 579 obtain a good estimate of the period of the direct phase, we can perform the origin time calibration 580 easily and efficiently. If the direct phases have high SNR and coda waves are partly coherent, 581 we may see a small flat top around Tmax or the maximum coherency value appears around the 582 theoretical maximally coherent time (as shown in Figure 24 the dashed line). Thus in this situation, 583 the estimate of the origin time will be affected and shows a small deviation. However, according to 584 our experience of processing the Uturuncu dataset, after calibration we can have a good estimation 585 of source origin times. Finally, in the stacked coherency trace, waveform coherency will decrease 586 to background noise level at one coherent analysis time window after the maximally coherent time. 587 For the Uturuncu dataset, only a few stations are available for source location, which negatively 588 affects the MCM imaging. However compared to other migration-based methods, in spite of the 589 very sparse monitoring array, MCM still obtains more reliable and precise location results by the 590 use of multichannel waveform coherency. A dense array with wide aperture and azimuth coverage 591 will greatly improve the imaging quality, especially when a long analysis time window is used. 592 When the stations are widely spread, the traveltime differences to different stations will be large. 593 Thus the migration result of the MCM will be better and the influence of the continuous coda 594 waves can also be reduced. High frequency information in the recorded data is important for 595 improving the imaging resolution. For the Uturuncu dataset, because of a low sampling rate, the

⁵⁹⁷ highest effective frequency is limited to 25 Hz. The volcano-tectonic earthquakes often contain ⁵⁹⁸ high-frequency content above this cut-off frequency. The insufficient sampling of the waveform ⁵⁹⁹ data (as can be seen in the Figure 3(b)) has limited the imaging resolution and quality. Despite the ⁶⁰⁰ sparse recording array and the lack of high frequency content, the MCM still obtains reliable event ⁶⁰¹ locations. When possible, we recommend volcano monitoring arrays record at at least 100 Hz to ⁶⁰² facilitate future automatic volcano-tectonic event determination.

For natural earthquakes, strong coda waves are often observed in the seismograms. The strong 603 coda waves can have a significant influence on the envelope of the waveforms and can also seri-604 ously affect the event detection using approaches such as STA/LTA and kurtosis. Thus the location 605 performance of the envelope, STA/LTA and kurtosis migration methods will be negatively affected 606 by the coda waves. Only when the coda waves of different stations are long-lasting and coherent, 607 they will have a negative impact on the MCM migration. The continuous coherent coda waves 608 will make the MCM source imaging ambiguous. One way to deal with coherent coda waves is to 609 increase the analysis time window for the coherency calculation. By using a longer time window, 610 the whole waveform train can be included in the coherent analysis and the direct P- or S-waves 611 as well as the coda waves are utilized to image the source event. Thus in this way the coherency 612 of the coda waves can be fully utilized to improve the event location, however at the expense of 613 reducing the imaging resolution. If coda waves are incoherent, a short analysis time window is 614 suggested to improve waveform coherency value and imaging resolution. 615

For event location at the Uturuncu, although the same velocity model used in obtaining the 616 event catalogue is utilized for location here, the event locations of the waveform migration method 617 are different to the event location in the catalogue, especially in event depth. The discrepancy 618 may come from different types of information being used in the event location. For events in the 619 catalogue, only the arrival times of the direct P- and S-waves obtained by manual picking are 620 used in the event location. However for waveform migration methods, the recorded waveforms 621 from different stations are directly used to locate the seismic event. MCM automatically identify 622 the maximum coherent time according to recorded traces and the predicted phase arrival times 623 are thus slightly different from the manually picked arrival times. Regardless of velocity model, 624

the location result of the arrival time based methods will be affected by the accuracy of manual picking, especially for low magnitude events. The location result of migration based methods is mainly affected by the signal-to-noise ratio and medium heterogeneity, which influence the recorded waveforms.

In the record sections (Figures 6 and 8), the recorded direct P- and S-wave arrivals at some 629 stations do not show a good consistency with the theoretical arrival times. And despite most direct 630 P-wave arrivals corresponding very well to the theoretical arrival times, the recorded S-waves 631 often arrive earlier than the theoretical S-wave arrival times. This discrepancy likely comes from 632 the velocity model used in the event location. Here we just applied a layered velocity model with a 633 constant v_p/v_s ratio of 1.75. In reality, the subsurface can have strong lateral velocity heterogeneity 634 as well as varying v_p/v_s ratio. The S-wave velocity model obtained by ambient noise tomography 635 (Jay et al. 2012) reveals the velocity heterogeneity in the Uturuncu area. If the velocity model is 636 very rough, it is worthwhile to adopt a method which can simultaneously locate the source and 637 update the velocity model. In this way, we can improve the event location accuracy and obtain a 638 more precise velocity model at the same time. By adjusting both the event location and velocity 639 model iteratively, the location results can match the arrivals of seismic phases more precisely. 640 Nevertheless, this is beyond the scope of this study. 641

For general waveform location methods based on the stacking of characteristic functions, the 642 imaging resolution in different directions is highly dependent on the array distribution. More geo-643 phones in a certain spatial direction increases resolution in that direction. However, if one direction 644 is better sampled than the other directions, the imaging results will be dominated by the waveform 645 stacking in that direction. Thus the imaging resolution in other directions (especially in the per-646 pendicular direction) will be degraded (as can be seen in the comparison between the first and 647 second rows in Figure 18). If we want to achieve equal resolution in different directions when 648 locating the source, evenly distributed geophones are required. However, MCM utilizes the co-649 herency between all possible receiver pairs, therefore the information from different directions 650 can achieve a better balance improving the MCM locations compared to the other methods. For 651 an irregularly distributed monitoring array, assuming n_p stations have been deployed in the pre-652

dominant direction, whereas n_c stations ($n_p > n_c$) are deployed in the non-predominant direc-653 tion. The contribution from non-predominant direction to the whole migrated volume for MCM 654 $(2n_pn_c/[(n_p+n_c)(n_p+n_c-1)])$ is always higher than that for conventional migration-based 655 methods $(n_c/(n_p + n_c))$. For MCM, due to the use of multichannel waveform coherency across 656 all the stations, the effective information from non-predominant direction can occupy a higher 657 proportion in migration compared to other conventional single-channel-based location methods. 658 Therefore, the imaging results of MCM are less affected by the irregular distribution of the re-659 ceivers, and the imaging resolution in different directions are well balanced. 660

As shown in the imaging results of the Aquistore real noise data, the location results of the 661 envelope and STA/LTA methods often show large deviations in depth. This is because the charac-662 teristic functions such as envelope and STA/LTA cannot represent the arrival times of the P- and 663 S-phases accurately. For envelope and STA/LTA, the maximum value of the characteristic func-664 tion often appears later than the correct arrival times of the P- and/or S-waves. For example, if the 665 source time function is a Ricker wavelet, the maximum value of the envelope is located at the peak 666 amplitude of the P- and S-phases, not at the accurate arrival times of the P- and S-phases, i.e. a half-667 period later. The characteristic function represents a transformation on the original waveform, and 668 the transformation on recorded waveforms of different stations can have different effects because 669 of noise, source radiation pattern, instrument response, etc. Thus the delayed times corresponding 670 to the correct arrival times can be different for different traces. This will lead to a trade-off between 671 the location depth and the origin time of the event. Finally, both the depth and origin time of the 672 location results can be biased. For the kurtosis method, due to the application of the derivative 673 of kurtosis (Langet et al. 2014), it can represent the arrival times of the P- and S-waves more ac-674 curately. Thus less deviations in depth are observed in the location results. However, the kurtosis 675 method is more affected by noise and irregular array geometry. The location results are not stable 676 compared to the other methods, and deviations in depth can easily appear when it cannot represent 677 the arrival time correctly. For MCM, due to the use of multichannel waveform coherency among 678 different traces, the maximum coherency value will appear at the correct arrival times of the P-679 and S-waves and waveform coherency will decrease rapidly when they deviate from the correct 680

source location (Shi et al. 2018b). There is less trade-off between location depth and origin time.
 Therefore, the MCM can accurately identify the source location and also the origin times with
 higher resolution.

Continuous coherent noise such as drilling noise remains a challenge for MCM. The coherent 684 noise which is continuous both in space and time will lead to high coherency values between 685 all receiver pairs, thus contributing to the failure of MCM when the coherent noise level is too 686 high. Removing the continuous coherent noise is key to overcoming this problem. If the coherent 687 noise in the recorded data falls into a specific frequency band, we can use frequency filtering or 688 frequency-wavenumber filtering to remove the coherent noise and improve the imaging quality. 689 For microseismic monitoring, the main coherent noise such as the drilling noise and injection 690 noise are often low frequency noise (less than tens of Hz), while the dominant frequency of the 691 microseismic signals are often relatively very high (from tens to thousands of Hz). Therefore, 692 this kind of low frequency noise can be separated and removed from the microseismic dataset by 693 filtering. Automatic quality control techniques such as weighting and filtering are effective ways 694 to mitigate the effects of noise and improve imaging quality. 695

696 6 CONCLUSIONS

In this paper, we applied the MCM method (Shi et al. 2018b) to locate microseismic events in 697 a reservoir and a volcanic setting in the presence of realistic noise. The location results of trig-698 gered volcano-tectonic earthquakes demonstrate the feasibility of using MCM method to locate 699 natural earthquakes recorded by sparse arrays. The MCM can automatically locate many triggered 700 events which are difficult and time consuming to manually pick. The MCM has the ability to lo-701 cate microseismic events which are otherwise often neglected by researchers. Using MCM, we can 702 efficiently obtain a more complete catalogue, which can help us better understand the subsurface 703 earthquake process. The newly obtained seismic catalogue at Uturuncu using MCM can be found 704 in the supplementary material. The predicted arrival times of P- and S-phases at different stations 705 are also attached, which can be used for further studies such as relocation. Compared to other 706 migration based methods, MCM shows more reliable location results and performs better in high 707

noise, sparse monitoring array and strong coda situations. The Aquistore real noise case demonstrates the excellent imaging performance of the MCM in the presence of strong realistic noise. Even though strong coherent noise exists in all traces, the MCM can still locate the source accurately. Usual quality control techniques such as the frequency filtering and weighting are feasible ways to remove coherent drilling or injection noise, the latter of which we employ in an automatic way. Compared to the other methods, the location results of the MCM have higher resolution and are more stable.

Computational efficiency tests of the MCM show that the MCM is highly scalable and par-715 allelizable. The parallel MCM code can achieve a high speedup ratio easily, which gives MCM 716 the ability to perform real time processing. Seismic location with sparse and/or irregularly dis-717 tributed monitoring array is problematic and difficult. MCM can expand the effective informa-718 tion used for locating by calculating multichannel waveform coherency across different stations, 719 thus in this way improving the location performance with sparse array. When the monitoring ar-720 ray is irregularly distributed, MCM imaging resolution in different directions can also be well 721 balanced due to the use of pairwise handling among all available stations. Compared to other 722 single-channel-based location methods, the location result of MCM is less affected by the irreg-723 ular and/or sparse distribution of the receivers, and the imaging resolutions in different direc-724 tions are higher and well balanced. The MCM code is open source and can be downloaded from 725 https://github.com/speedshi/seisloc.The MCM code is written in FORTRAN and further 726 developments of the MCM software will be released in the future. 727

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Figure 1. Computational efficiency analysis of the MCM location. The efficiency test is performed on the Intel E5-2670(2.6GHz) processor. (a) The computational times for different numbers of imaging points and searched origin times. Black line and the bottom X-axis show the variation of computational times with the number of imaging points, when the number of origin times and stations are fixed as 1000 and 40. Blue line and the top X-axis show the variation of computational times with the number of origin times, when the number of imaging points and stations are fixed as 100000 and 40. (b) The computational times for different numbers of stations, when the number of origin times and imaging points are fixed as 1000 and 100000. Black line and the bottom X-axis show the variation of computational time with the number of stations. Blue line and the top X-axis show the variation of computational time with the number of unique station pairs. Program runs on one core for (a) and (b). (c) The computational times (black line and left Y-axis) and speedup ratios (blue line and right Y-axis) when different numbers of cores are used. Red dashed line show the theoretical speedup ratios. The number of origin times, imaging points and stations are fixed as 1000, 100000 and 40. (d) The required cores used for real time processing under different numbers of stations and imaging points. Different black lines show the scenarios for different numbers of imaging points. The red dot shows the scenario for the following Uturuncu dataset, where 14 stations are deployed, 283556 imaging points are scanned and 68 cores are required for real time processing.



Figure 2. Location of the seismic stations and Uturuncu volcano (UTM zone: 19K). The stations are represented by gray triangles. Two local volcano-tectonic earthquakes in the catalogue are represented by red stars. The Uturuncu is located in the middle of the figure. The color in the figure represents elevation relative to the sea level. The lower left part exhibits a regional map, in which the red rectangle shows the research area. The lower right part exhibits the velocity model used in the event location, in which the red and blue lines show the P- and S-wave velocities. The white rectangle shows the imaging area (shown as Figure 11) for the four hours of continuous data.



Figure 3. The recorded three component waveforms at station UTCA for the two shallow, local volcanotectonic earthquake. The blue and red lines show the arrivals of P- and S-waves respectively. (a) Waveforms for the first event. The instrument response has been removed and the waveforms are filtered using a bandpass filter of 5-23 Hz. (b) Waveforms for the second event. The instrument response has been removed and the waveforms are filtered using a bandpass filter of 5-21 Hz.



Figure 4. Migration profiles through the maximum migrated value for the first volcano-tectonic earthquake. The dark stars show the corresponding seismic event in the catalogue obtained by manual picking. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows YZ profiles, second row shows XZ profiles, third row shows XY profiles.



Figure 5. The stacking functions of the MCM method at the position of the maximum migrated value. The red line shows the origin time of the event in the catalogue obtained by manual picking. (a) The stacking function for event 1. The time is relative to 2009-05-25 02:00:00. (b) The stacking function for event 2. The time is relative to 2009-05-08 04:30:00.

Table 1. Location results of different waveform migration methods for the Uturuncu shallow volcanotectonic earthquake and comparison with the event in the catalogue. The origin time is relative to 2009-05-25 02:00:00 (UTC).

	Event location				Deviation from manual traveltime location			
	X (km)	Y (km)	Z (km)	T_0 (s)	$\Delta X(m)$	$\Delta Y\left(\mathrm{m} ight)$	$\Delta Z(m)$	ΔT_0 (s)
Catalogue	7530.316	680.543	-0.269	1708.2	-	-	-	-
Envelope	7533.4	682.2	-3.2	1709.8	3084	1657	2931	1.6
STA/LTA	7528.0	682.6	1.0	1709.4	2316	2057	1269	1.2
Kurtosis	7530.1	679.7	2.0	1708.4	216	843	2269	0.2
Coherency	7530.9	681.1	0.2	1708.3	584	557	469	0.1


Figure 6. The three component record sections of the first event. The predicted P- and S-wave arrival times for this event in the catalogue and the event located by MCM are marked by solid and dashed lines respectively. The blue and red colors show the arrival times of the direct P- and S-wave respectively. The time in the figure is relative to 2009-05-25 02:00:00 (UTC).



Figure 7. Migration profiles through the maximum migrated value for the second volcano-tectonic earthquake. The dark stars show the corresponding seismic event in the catalogue obtained by manual picking. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows YZ profiles, second row shows XZ profiles, third row shows XY profiles.

Table 2. Location results of different waveform migration methods for the Uturuncu shallow volcanotectonic earthquake and comparison with the event in the catalogue. The origin time is relative to 2009-05-08 04:30:00 (UTC).

	Event location				Deviation from manual traveltime location			
_	X (km)	Y (km)	Z (km)	T_0 (s)	$\Delta X(m)$	$\Delta Y\left(\mathbf{m} ight)$	$\Delta Z(m)$	ΔT_0 (s)
Catalogue	7540.866	687.419	-1.523	1581.3	-	-	-	-
Envelope	7540.9	690.5	-1.6	1582.2	34	3081	77	0.9
STA/LTA	7541.6	686.6	-3.6	1582.9	734	819	2077	1.6
Kurtosis	7541.4	687.4	0.2	1580.9	534	19	1723	0.4
Coherency	7540.7	687.6	0.4	1580.7	166	181	1923	0.6



Figure 8. The three component record sections of the second event. The predicted P- and S-wave arrival times for this event in the catalogue and the event located by MCM are marked by solid and dashed lines respectively. The blue and red colors show the arrival times of the direct P- and S-wave respectively. The time in the figure is relative to 2009-05-08 04:30:00 (UTC).



Figure 9. The recorded Z component waveforms at station UTCA. The recording time ranges from 06:00:00 to 10:00:00 (UTC). The waveforms within the blue rectangle are enlarged. The instrument response has been removed and the waveforms are filtered using a bandpass filter of 4.2-21.6 Hz.



Figure 10. The maximum coherency value at each searched origin time for the four hours of continuous data. The time interval is 0.08 s. The part in the blue rectangle is enlarged. The red points show the 322 verified seismic events.



Figure 11. The earthquake locations on the horizontal and vertical profiles. Black dots show the 114 event locations in the existing catalogue. The color-coded dots show the verified 322 event locations for the MCM. The color represents the origin times of earthquake events.



Figure 12. Horizontal and vertical profiles at the maximum value of the migration volume for seismic event 1. Color represents the migration value (coherency). Black star represents the event location in the catalogue.

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Figure 13. Three component record sections for seismic event 1. The predicted P- and S-wave arrival times are marked by blue and red crosses, respectively. Left panel: record sections for the MCM location result. Right panel: record sections for the catalogue location result. The time in the figure is relative to 2010-02-27 06:00:00 (UTC).



Figure 14. Horizontal and vertical migration profiles and three component record sections for seismic event 2, which is newly detected by MCM and not in the existing catalogue. The predicted P- and S-wave arrival times are marked by blue and red crosses on the record sections, respectively. Left panel: horizontal and vertical profiles at the maximum value of the migration volume. Color represents the migration value and black star shows the final event location. Right panel: record sections for this event.



Figure 15. Histogram of Uturuncu triggered events from the M_w 8.8 Maule earthquake for the four hours (6 am to 10 am) in 5-min bins. There are no seismic events from 6:00 am to 06:40 am. Blue dashed lines show the approximate arrival time of surface wave trains. G1/R1 represents the minor-arc Love (G1) and Rayleigh (R1) waves. X2/X3, G2/G3 and R2/R3 represent different surface wave overtones (Jay et al. 2012).



Figure 16. (a) Aquistore permanent seismic array geometry. Geophones are denoted by red dots alongside the station number, while the observation and injection wells are illustrated by yellow triangles. (From Birnie et al. (2016)). (b) P- and S-wave velocity model and density model in Aquistore area. The red color highlights two target layers where the seismic events are located. (c) The numerical model space of the Aquistore area. Vertical (d) XZ and (e) YZ profiles of the numerical model. The red stars shows the locations of two seismic events, whose depth are 1.85 km and 3.25 km respectively. Blue points represent the surface geophones. The yellow color exhibits the imaging area of the shallow event. Source radiation patterns are shown in the vertical profiles using a beach ball with red and blue colors.



Figure 17. (a) The recorded waveform data at station 30 before (top) and after (bottom) adding real noise. The blue and red crosses show the arrivals of P- and S-phases. (b) The synthetic noise-free seismic profile. (c) The seismic profile after adding real noise. The SNR is 0.5.



Figure 18. Profiles of the migration results through the true source location for the four methods. The SNR is 1. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows YZ profiles, second row shows XZ profiles, third row shows XY profiles.



Figure 19. The absolute amplitude ratios for different stations under different SNR scenarios. The absolute amplitude ratio of different traces is defined as the ratio of the average absolute amplitudes of a trace to the average absolute amplitude of all traces. The black dashed line shows an absolute amplitude ratio of 1.5.

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Figure 20. Profiles of the migration results through the true source location with automatic quality control scheme (weighting and filtering). The SNR is 0.025. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows YZ profiles, second row shows XZ profiles, third row shows XY profiles.



Figure 21. The location errors of the four methods under different SNRs with (a) original data and (b) automatic quality control scheme (weighting and filtering).

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Figure 22. Migration profiles through the true source location of the deep event. The SNR is 1. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows YZ profiles, second row shows XZ profiles, third row shows XY profiles.



Figure 23. The stacking functions of the four methods at the true source location of the deep event for the Aquistore noise data. The red dashed lines show the origin time of the source time function and the black dashed lines show the end time of the source time function. The pink areas around -0.6 s and 0.7 s highlight the time range where P-/S-phases move into the stacking window of the S-/P-phases when searching for origin time. The SNR is 1.



Figure 24. Schematic diagram showing recorded waveforms and the corresponding stacked coherency trace. Tw is the length of coherent analysis time window, and T is the period of direct wave. The orange dot shows the arrival time of direct wave, and the red dot shows the maximum coherency value at the stacked coherency trace. For the stacked coherency trace, the solid line shows the maximum coherency value appearing at T time after the rise of waveform coherency, and the dashed line shows the maximum coherency value appearing within T time after the rise of waveform coherency.



Supplementary Materials

Figure S.1. Vertical profiles (YZ profiles) through the true source location of the migration results under different SNRs for the four methods. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows the results when data is free of noise, second row for SNR is 1, third row for SNR is 0.5, fourth row for SNR is 0.25, fifth row for SNR is 0.025.



Figure S.2. Vertical profiles (XZ profiles) through the true source location of the migration results under different SNRs for the four methods. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows the results when data is free of noise, second row for SNR is 1, third row for SNR is 0.5, fourth row for SNR is 0.25, fifth row for SNR is 0.025.



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Figure S.3. Horizontal profiles (XY profiles) through the true source location of the migration results under different SNRs for the four methods. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows the results when data is free of noise, second row for SNR is 1, third row for SNR is 0.5, fourth row for SNR is 0.25, fifth row for SNR is 0.025.





Figure S.4. Vertical profiles (YZ profiles) through the true source location of the migration results with automatic quality control scheme (weighting and filtering) under different SNRs. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows the results when data is free of noise, second row for SNR is 1, third row for SNR is 0.5, fourth row for SNR is 0.25, fifth row for SNR is 0.025.



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Figure S.5. Vertical profiles (XZ profiles) through the true source location of the migration results with automatic quality control scheme (weighting and filtering) under different SNRs. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows the results when data is free of noise, second row for SNR is 1, third row for SNR is 0.5, fourth row for SNR is 0.25, fifth row for SNR is 0.025.



Figure S.6. Horizontal profiles (XY profiles) through the true source location of the migration results with automatic quality control scheme (weighting and filtering) under different SNRs. The dark star in the center shows the true source location. The first column shows results of envelope, second column for STA/LTA, third column for kurtosis, fourth column for MCM. The first row shows the results when data is free of noise, second row for SNR is 1, third row for SNR is 0.5, fourth row for SNR is 0.25, fifth row for SNR is 0.025.