

1 **Lower Crustal Seismicity on the Eastern Border Faults of the Main Ethiopian Rift**

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3 **Enter authors here: Sacha Lapins<sup>1</sup>, J-Michael Kendall<sup>2</sup>, Atalay Ayele<sup>3</sup>, Matthew Wilks<sup>1</sup>,**  
4 **Andy Nowacki<sup>4</sup>, and Katharine V Cashman<sup>1</sup>.**

5 <sup>1</sup>School of Earth Sciences, University of Bristol, UK.

6 <sup>2</sup>Department of Earth Sciences, University of Oxford, UK.

7 <sup>3</sup>Institute of Geophysics, Space Science and Astronomy, Addis Ababa University, Ethiopia.

8 <sup>4</sup>School of Earth and Environment, University of Leeds, UK.

9

10 Corresponding author: Sacha Lapins (sacha.lapins@bristol.ac.uk)

11

12 **Key Points:**

- 13 • A sequence of lower crustal seismicity is identified 30km east of Corbetti caldera on the  
14 eastern border faults of the Main Ethiopian Rift
- 15 • Mixture distributions are used to overcome misinterpretations from large uncertainty in  
16 individual event locations
- 17 • Spatial and temporal characteristics indicate that the source of this seismicity relates to  
18 fluid or magmatic processes

## 19 **Abstract**

20 Lower crustal seismicity is commonly observed in continental rift zones despite the crust at such depths  
21 being ductile enough to prohibit brittle failure. The source of such deep seismicity across the East African Rift  
22 remains an outstanding question. Here we present analysis of an isolated cluster of lower crustal earthquakes located  
23 on the eastern border faults of the Main Ethiopian Rift, near the Corbetti caldera. Lower crustal earthquakes have  
24 not previously been observed in this area. Phase arrival times were determined using an automated picking approach  
25 based on continuous wavelet transform and statistical changepoint detection methods. We overcome  
26 misinterpretations from large hypocentre depth errors by considering mixture distributions for all events and their  
27 associated uncertainties. These mixture distributions represent probability density functions of any event occurring  
28 at a given depth. The mixture distribution mode for a variety of different velocity models and error parameters  
29 remained stable at a depth of 28 – 32 km, with the vast majority of maximum likelihood estimates for individual  
30 hypocenters located at depths of 25 – 35 km. Most events occur over a two-month period, with 90% of cumulative  
31 seismic moment occurring during March and April 2012. The ephemeral and localized nature of this seismicity,  
32 combined with low event magnitudes and regional hydrothermal / magmatic activity, all suggest that these lower  
33 crustal events are likely related to fluid or magmatic processes. Plausible mechanisms include the movement of  
34 magma and/or exsolution of volatiles at depth causing transient high strain rates and pore fluid pressures that induce  
35 seismicity.

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## 37 **Plain Language Summary**

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## 40 **1 Introduction**

41

42 Seismicity associated with continental rift zones can reach depths of more than 30 km (e.g., Albaric et al.,  
43 2014; Keir et al., 2009; Reyners et al., 2007; Yang & Chen, 2010) despite the general conception that the crust at  
44 such depths is too ductile for brittle failure to occur (Chen & Molnar, 1983). In the East African Rift (EAR; Fig 1),  
45 deep crustal, and even upper mantle, seismicity has been observed along the younger eastern and western branches  
46 of the EAR (Albaric et al., 2014; Lavayssière, Drooff, et al., 2019; Lindenfeld & Rumpker, 2011; Nyblade &  
47 Langston, 1995; Yang & Chen, 2010), as well as beneath the western margins of the more well-developed,  
48 magmatic Main Ethiopian Rift (MER) at the northern end of the of the EAR (Keir et al., 2009; Fig 1). One  
49 explanation is that these earthquakes are caused by slip on border faults into relatively cold or strengthened lower  
50 crust (Doser & Yarwood, 1994; Jackson & Blenkinsop, 1993; Zhao et al., 1997). Such an explanation requires the  
51 underlying mantle to be cool and/or strong enough for earthquakes to occur, for which there may be compelling  
52 evidence beneath less magmatic segments on the western branch of the EAR (Lindenfeld & Rumpker, 2011; Yang  
53 & Chen, 2010). However, beneath the considerably more magmatic MER, where the lithospheric mantle has high

54 enough temperatures to sustain partial melt (Kendall et al., 2005; Hammond et al., 2014), the lower crust and upper  
55 mantle are deemed too ductile for brittle failure; here deep seismicity is thought to be induced by magmatic  
56 processes and fluids (Keir et al., 2009; Reyners et al., 2007; Soosalu et al., 2010). In other oceanic and continental  
57 rift settings, such as Iceland and the Taupo rift in New Zealand, deep seismicity has been associated with high strain  
58 rates / pore pressures (e.g. Greenfield & White, 2015; Soosalu et al., 2010; White et al., 2011), magma/fluid  
59 emplacement (Smith et al., 2016) and the weakening of border faults (e.g., Reyners et al., 2007) arising from melt  
60 movement and/or the exsolution of volatiles. However, the mechanism for deep seismicity in the EAR, where the  
61 level of magmatism varies from section to section, is still debated (e.g., Weinstein et al., 2017).

62  
63 In this paper, we examine an isolated cluster of lower-crustal seismicity approximately 30 km east of the  
64 Corbetti caldera on the eastern border faults of the MER, where no previous lower crustal seismicity has been  
65 observed. The distribution mode of this seismicity lies at a depth of 28 – 32 km, directly beneath the Wondo Genet  
66 scarp (Fig 2). This scarp is thought to relate to an intersection between the rift border and the easternmost end of a  
67 cross-cutting fault structure that pre-dates the rift (Lloyd, Biggs, Wilks, et al., 2018; Fig 2), with the interaction of  
68 these two faults potentially influencing seismicity in the area. We determine event phase-arrival times with high  
69 precision using wavelet transform and statistical changepoint detection methods. We attempt to reduce the  
70 variability of individual event locations, largely caused by poor seismic array coverage relative to the region of  
71 seismicity, by considering the mixture distribution of all events observed for a given velocity model and set of  
72 location algorithm parameters. The temporal and spatial distribution of these events suggests that these isolated  
73 events may be caused by a single, ephemeral intrusion at lower crustal depths that causes overpressure, inducing  
74 seismicity, or hot fluids reducing the effective normal stress on the faults.

75

## 76 **2 The Main Ethiopian Rift (MER) and Corbetti Caldera**

77

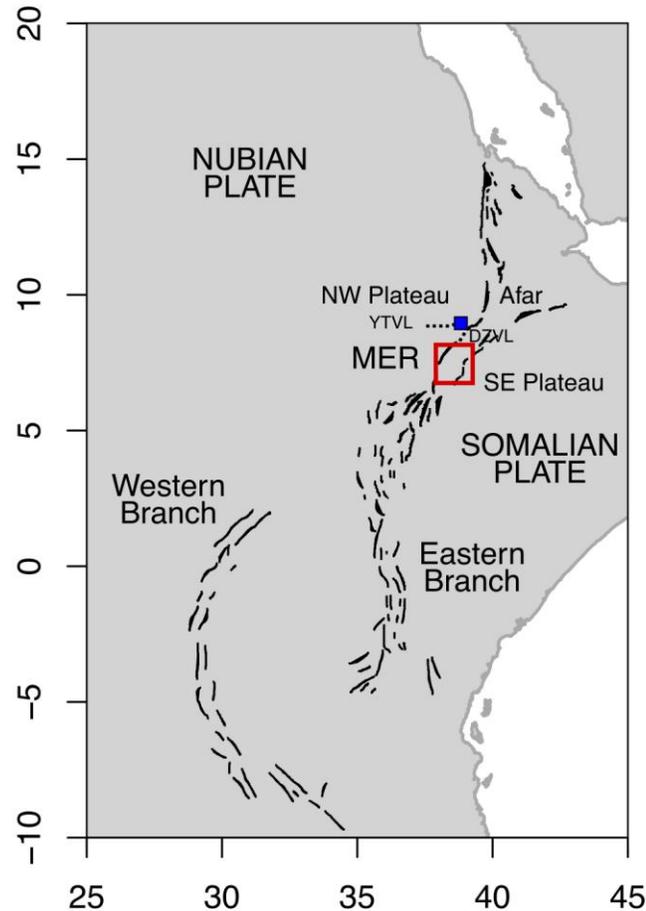
78 The MER is an area of continental rifting that forms the major divergent plate boundary between the  
79 Nubian and Somalian tectonic plates at the northern end of the EAR (Fig 1). It is thought to have initiated and  
80 developed asynchronously along its length (e.g., Wolfenden et al., 2004), with the development of different sectors  
81 influencing magmatism, strain and crustal thickness across the region (Keir et al., 2015; Muluneh et al., 2017). Most  
82 regional seismicity occurs along the axis of the MER (Wilks, 2016), where the seismogenic layer is constrained to  
83 upper crustal depths ( $< \sim 16$  km; Muluneh et al., 2018). However, earthquakes as deep as  $\sim 35$  km (the approximate  
84 thickness of the crust in this region; Dugda et al., 2005; Ebinger et al., 2017; Stuart et al., 2006) have been identified  
85 on the western margin of the MER, associated with areas of recent volcanism along the Debre-Zeit and Yerer-Tullu  
86 Wellel Volcanotectonic Lineaments near Addis Ababa (Keir et al., 2009). Additional lower crustal seismicity at  
87 depths of  $\sim 22$  km, which is still below the Brittle-Ductile Transition (BDT) zone for the region (16km depth;  
88 Muluneh et al., 2018), has been observed beneath the rift-adjacent north-western (NW) Ethiopian Plateau (Keir et  
89 al., 2009), north of Addis Ababa. The location of this seismicity in volcanic areas supports the idea of magma

90 emplacement below/into the lower crust as a source of lower crustal seismicity in the MER (e.g., Keir et al., 2009;  
 91 Soosalu et al., 2010), rather than fault slip within a relatively cold or strengthened lower crust.

92

93 The Corbetti caldera is the southernmost silicic center along the rift. It lies toward the eastern side of the  
 94 MER and formed at  $182 \pm 18$  ka (Lloyd, Biggs, Wilks, et al., 2018). Its magmatic and hydrothermal processes are  
 95 thought to be influenced by a cross-cutting fault structure that pre-dates the rift and may extend as far as the rift  
 96 border,  $\sim 30$  km to the east (Lloyd, Biggs, Wilks, et al., 2018). Seismicity beneath the caldera appears to be  
 97 constrained to the uppermost crust (Lavayssière, Greenfield, et al., 2019; Lloyd, Biggs, Birhanu, et al., 2018),  
 98 although previous seismic monitoring in the area has been limited.

99 Prior to the events analyzed in this paper, no other sequence of lower crustal seismicity had been observed  
 100 beneath the rift itself, the Corbetti caldera, the eastern border faults of the MER, or the seismically quiet south-  
 101 eastern (SE) Somalian plateau to the east.



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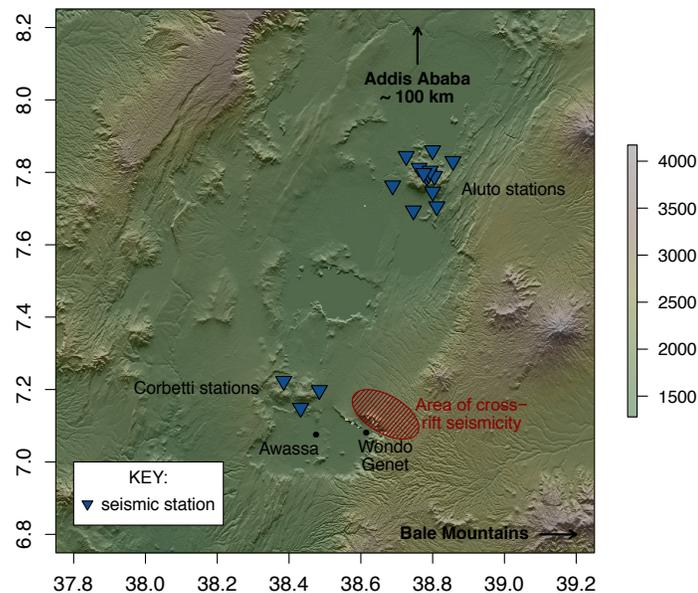
103 *Figure 1 – Map of East African Rift (EAR) with schematic rift-bounding faults after Foster and Jackson (1998).*  
 104 *Horizontal and vertical axes show longitude and latitude, respectively. Blue square is Addis Ababa. Red box is*

105 *location of map in Figure 2. Previous lower crustal seismicity has been observed along less magmatic sections of*  
 106 *the western and eastern branches of the EAR to the south, as well as beneath the Debre-Zeit and Yerer-Tullu Wellel*  
 107 *Volcanotectonic Lineaments (DZVL, dotted line southeast of Addis Ababa, and YTVL, dotted line west of Addis*  
 108 *Ababa, respectively) on the western margins of the Main Ethiopian Rift (MER).*

109

### 110 **3 Seismic Array Geometry**

111 Two local seismic arrays were deployed at Aluto (Wilks et al., 2017) and Corbetti (Wilks, 2016) volcanoes  
 112 throughout 2012 and 2013, with 12 and 7 broadband seismometers, respectively, deployed at each volcano, although  
 113 only subsets of around 15 of these instruments were operational at any given time throughout the study period (Fig  
 114 2). The relative location of these two local arrays to the observed seismicity leaves a substantial array gap (~ 270  
 115 degrees), with no other regional seismic networks operating at the time.



116

117 *Figure 2 – Map of available instruments from Corbetti and Aluto volcano seismic arrays, with general area of*  
 118 *analyzed seismicity highlighted around 30 km to the east of Corbetti. Map color denotes surface altitude in meters*  
 119 *above sea level. Horizontal and vertical axes show longitude and latitude, respectively.*

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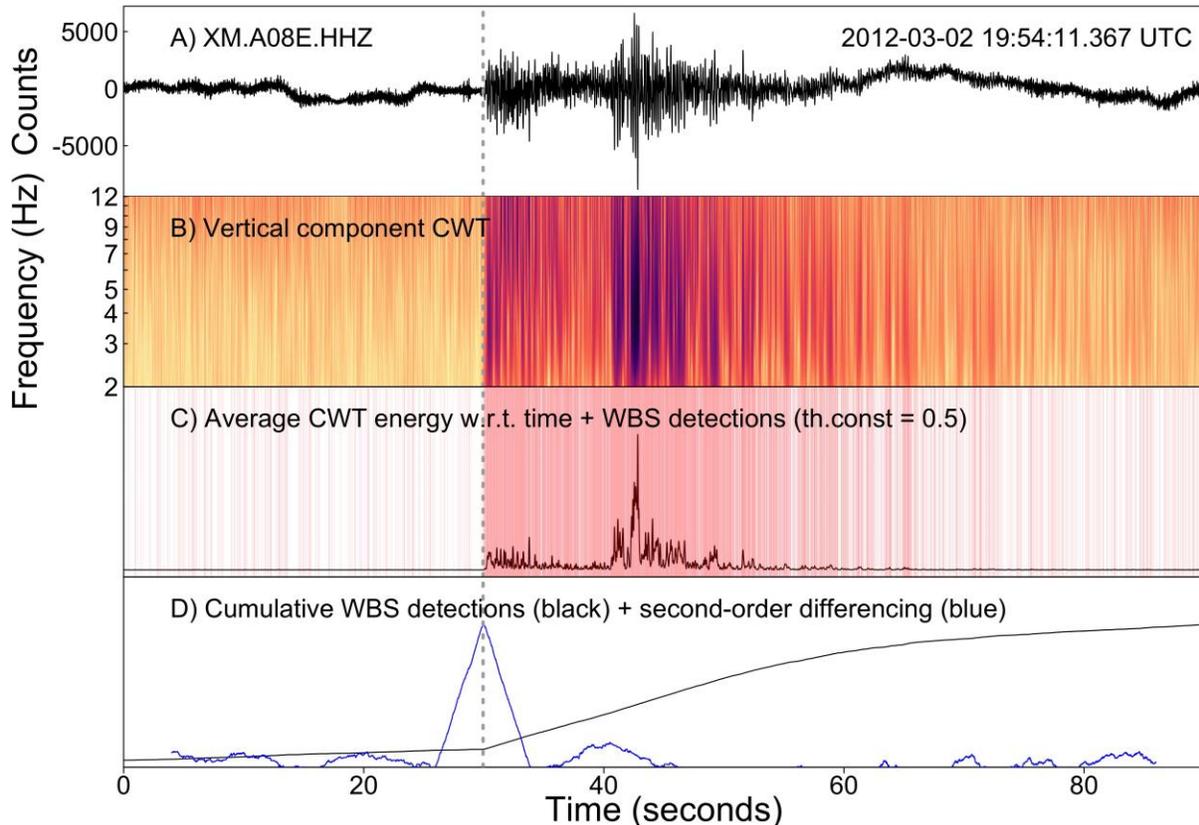
### 121 **4 Phase Arrival Picking**

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123 Phase arrival times were identified manually across the full deployment ( $> 3000$  events over 2 years  
124 continuous data; Wilks, 2016). The anomalous subset of apparent lower crustal seismicity was identified by  
125 subsequent event location (number of events = 134). Initial phase arrival picks for this subset had large pick errors  
126 and large predicted travel-time (TT) residuals, sometimes on the order of several seconds; for this reason, the first  
127 step in this study was to reduce this source of error in hypocentre estimation through a time-frequency-based  
128 automated picking approach (Figs 3 and 4). This approach was based on continuous wavelet transform (CWT)  
129 spectral analysis (e.g., Lapins et al., 2020) and statistical changepoint detection (Fryzlewicz, 2014).

130 P-wave arrivals were characterized by producing CWT scalograms between 2 and 12 Hz (e.g., Lapins et al.,  
131 2020; Fig 3b) for the raw vertical component traces at each station (Fig 3a). We then determined the average  
132 wavelet energy between these frequency bounds for each time sample point (black line in Fig 3c) using both low-  
133 and standard-frequency Morlet wavelets (Morlet central frequency parameter  $\omega_0 = 1$  and  $\omega_0 = 6$ , respectively).  
134 Changes in average wavelet energy with time were detected using Wild Binary Segmentation (WBS; Fryzlewicz,  
135 2014), an *a posteriori* changepoint detection method that segments the signal by a given statistical property, which  
136 in this case is mean energy. The WBS ‘threshold’ parameter for determining a change in mean energy was set very  
137 low (threshold constant  $c = 0.5$ ), which led to a near-uniform rate of ‘false positive’ changepoint detections  
138 during pre-event noise (1 – 2 detections per second) and a much higher rate of changepoint detections at the P-wave  
139 arrival and during the coda ( $>> 1$  detection per second; vertical red lines in Fig 3c). The cumulative number of  
140 detected changepoints within a signal window (90 s) was then used to identify P-wave arrival through second-order  
141 differencing (lag = 4 s) and a simple trigger algorithm (Fig 3d).

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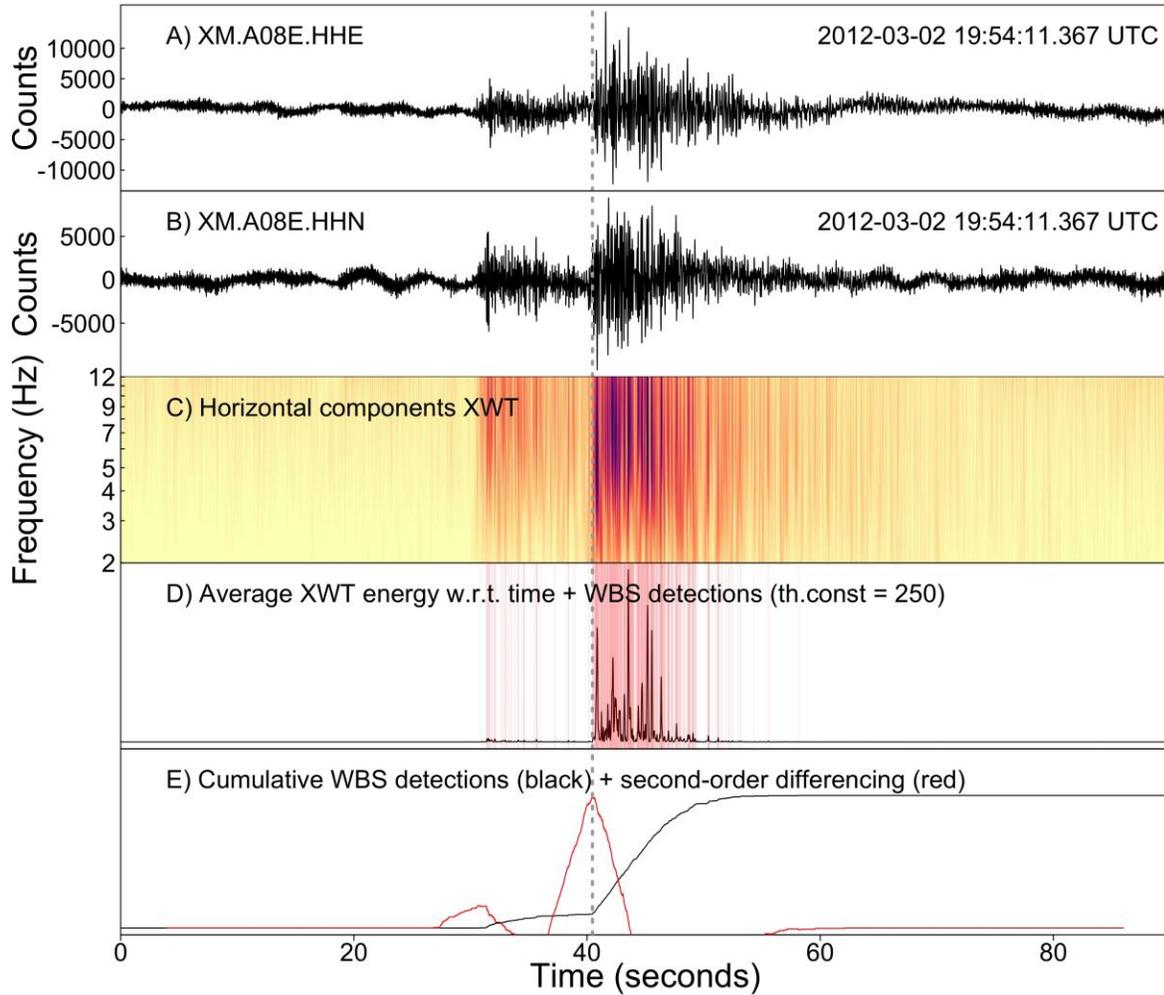


143

144 *Figure 3 – A) Raw vertical component trace; B) CWT scalogram between 2 and 12 Hz for trace above (Morlet*  
 145 *wavelet with central frequency  $\omega_0 = 1$ ); C) Average wavelet energy (black line) with WBS change point*  
 146 *detections marked with semi-transparent red lines; D) Cumulative number of WBS detections (black line) with second-order*  
 147 *differencing (lag = 4 s; blue line) – sharp blue peak and dashed grey line indicate determined P-wave arrival time.*

148 S-wave arrivals were identified in a similar manner to P-wave arrivals but using the cross-wavelet  
 149 transform (XWT) derived from the two horizontal component CWT scalograms at each station (Fig 4). This had the  
 150 effect of reducing incoherent background noise while enhancing coherent signal across the two horizontal  
 151 components (Fig 4b). Again, the average wavelet energy was determined at each time sample point for low- and  
 152 standard-frequency Morlet wavelets. Finally, WBS and second-order differencing was used to identify S-wave  
 153 arrivals common to both horizontal components (Figs 4c and d), this time with a greater WBS threshold constant  
 154 ( $c = 250$  for Aluto stations,  $c = 5000$  for nearby Corbetti stations) to avoid a higher rate of change point  
 155 detections around P-wave arrivals. We use different threshold constants across the two arrays due to higher  
 156 amplitude values in the raw signal at the nearer Corbetti stations.

157



158  
 159 *Figure 4 – A) Raw EW horizontal component trace; B) Raw NS horizontal component trace; C) Cross wavelet*  
 160 *transform (XWT) scalogram between 2 and 12 Hz for both horizontal components (Morlet wavelet with central*  
 161 *frequency  $\omega_0 = 1$ ); D) Average XWT energy (black line) + WBS changepoint detections marked with semi-*  
 162 *transparent red lines; E) Cumulative number of WBS detections (black line) and double-differenced cumulative*  
 163 *number of WBS detections (lag = 4 s; red line). Sharp red peak and dashed grey line indicate determined S-wave*  
 164 *arrival time.*

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All P- and S-wave picks were manually checked against raw and filtered traces, CWT scalogram images and additional STA/LTA (e.g., Withers et al., 1998) traces to assess quality and uncertainty. Arrival times for highly emergent signals were sometimes picked late by the CWT-WBS approach outlined above, likely a consequence of the lower frequency onset of these signals, which is outside of the frequency range encompassed by the CWT scalograms (2 – 12 Hz). These arrival times were adjusted manually using visual determination on wider frequency band CWT scalograms and raw traces. Arrival time picks that were difficult to confirm manually (e.g., in the presence of very low signal-noise ratio, SNR) were removed.

173 Picks made using the low-frequency Morlet wavelet ( $\omega_0 = 1$ ), which yields improved time resolution at  
 174 the cost of frequency resolution (Addison et al., 2002; Lapins et al., 2020), were more accurate for signals with  
 175 medium to high SNR, whilst picks based on the standard Morlet wavelet ( $\omega_0 = 6$ ) were more accurate for very  
 176 noisy signals because of greater frequency localization of pertinent signal features among ambient noise (Lapins et  
 177 al., 2020).

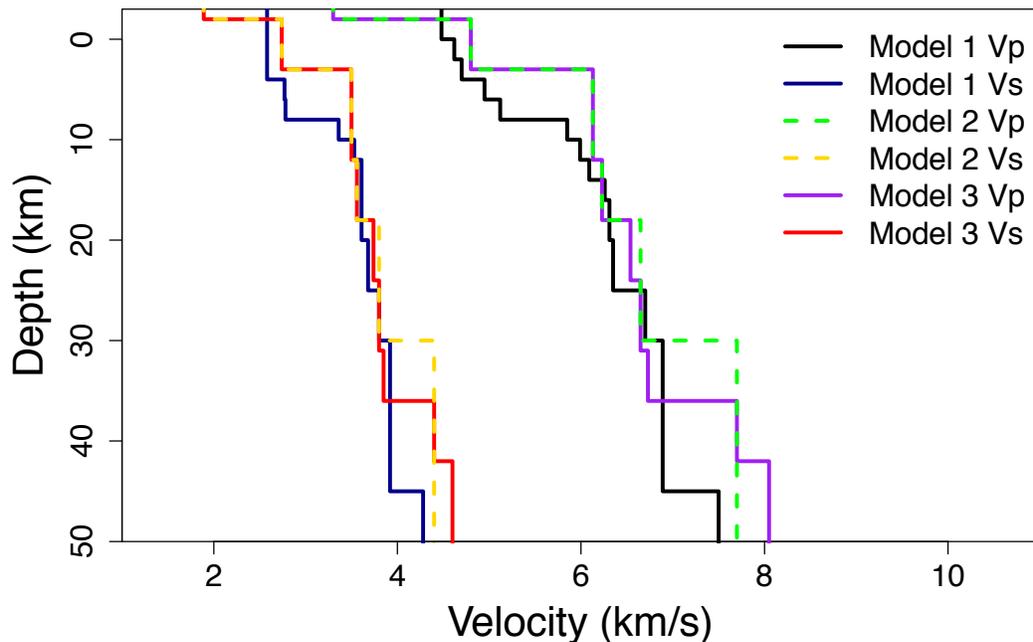
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## 179 **5 Absolute Event Locations**

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181 Events were located using NonLinLoc (Lomax et al., 2000) and three different regional velocity models  
 182 (Daly et al., 2008; Mackenzie et al., 2005; details in Fig 5 caption); travel-time error was varied to assess the  
 183 stability of event locations and determine potential bias in absolute hypocentre locations (e.g., artificially deep  
 184 locations). Only events with a total of seven or more phase arrival times, including at least one S phase, were located  
 185 (58 events in total). We include at least one S phase to improve spatial constraints on hypocentre locations (e.g.,  
 186 Lomax et al., 2009), which can have large uncertainties when using stations with a large array gap or that are far  
 187 from the event.

188



189

190 *Figure 5 – Velocity models used for locating events ( $V_p$  = P-wave velocity;  $V_s$  = S-wave velocity). Model 1 is from*  
 191 *Table 1 in Daly et al. (2008); Model 2 is approximated from Fig 5 in Daly et al. (2008); Model 3 is approximated*  
 192 *from Fig 6 in Mackenzie et al. (2005).  $V_p/V_s$  ratio of 1.75 (average from Table 1 in Daly et al., 2008) used for*  
 193 *Models 2 and 3 as only  $V_p$  given in original publications.*

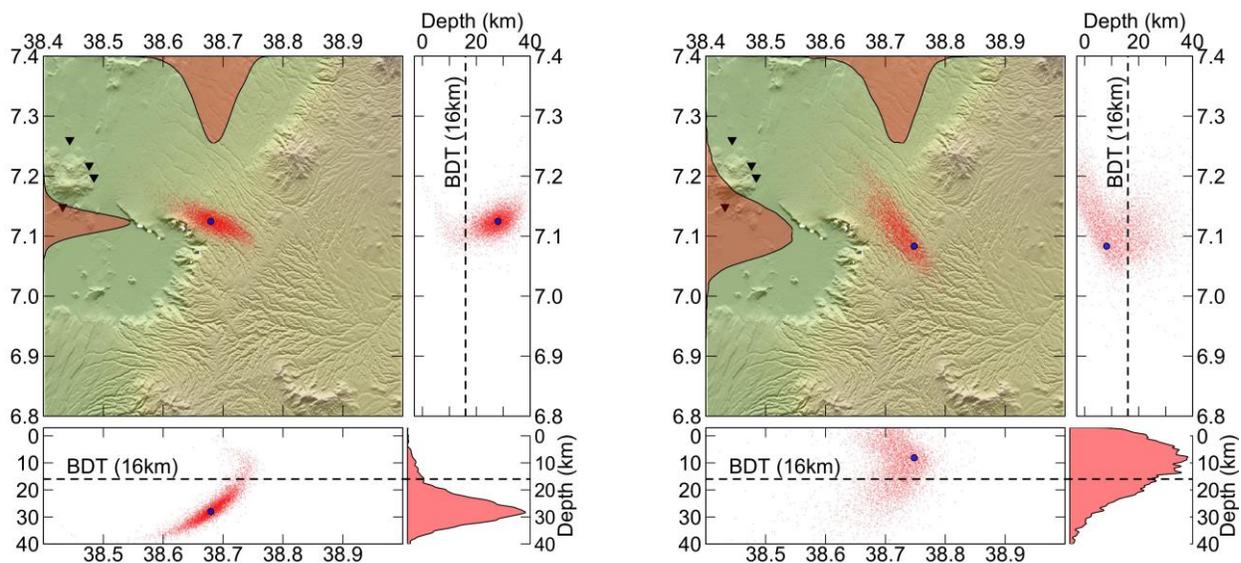
194

195 The significant array gap in this case (approximately 270 degrees) meant that individual event locations did  
 196 indeed have large error distributions, although deeper events were generally better constrained than shallower events  
 197 (Fig 6). For this reason, we also examine mixture distributions (a weighted combination of probability distributions  
 198 for each individual event location) for all events, with a given set of location parameters (Fig 7), to determine  
 199 stability and likelihood of event depth estimates using the whole population of events in the study. More formally, a  
 200 (finite) mixture distribution,  $f$ , is a convex combination of  $n$  component probability distributions,  $p_1(\mathbf{x}), \dots, p_n(\mathbf{x})$ ,  
 201

$$f(\mathbf{x}) = \sum_{i=1}^n w_i p_i(\mathbf{x}), \quad (1)$$

202 with weights  $w_1, \dots, w_n$  such that  $w_i \geq 0$  and  $\sum w_i = 1$ . Here, the component distributions,  $p_i(\mathbf{x})$ , are the  
 203 individual posterior probability density functions (PDFs) for each event's hypocentre location, estimated by  
 204 NonLinLoc, where the vector  $\mathbf{x}$  represents the three-dimensional spatial x, y, z hypocentre location and  $i$  is event  
 205 index  $1, \dots, n$ . A mixture distribution preserves the required properties of probability distributions (non-negativity  
 206 and integrating to 1) and is therefore itself a probability distribution.

207



208

209 *Figure 6 – Examples of individual event locations from NonLinLoc. Left: example of ‘well-constrained’ deep event –*  
 210 *depth error estimates (red scatter and density curves) are quite large, spanning ~ 20 km depth, although clearly*  
 211 *below the BDT and far more compact than those for shallow events. Right: an example of a ‘typical’ shallow event –*  
 212 *red scatter and density curves representing error estimates are more diffuse and multimodal, suggesting greater*  
 213 *uncertainty in event location. Red scatter and density curves are comprised of 5,000 samples drawn from the*  
 214 *posterior PDF for each event hypocentre location.*

215

216 We use mixture distributions to estimate the overall population distribution for all event locations and  
217 assess the likelihood of events occurring at a given depth (from here on we use the terms *overall population*  
218 *distribution* and *mixture distribution* interchangeably). In practice, probability distributions for individual earthquake  
219 locations are nonlinear and may be mathematically intractable. For this reason, earthquake location distributions are  
220 often estimated through probabilistic sampling of their complete posterior distribution (e.g., Gesret et al., 2015;  
221 Lomax et al., 2000). The number of samples used for estimating each component distribution is determined by its  
222 corresponding weight with regards to the mixture distribution (Eq. 1). As each component distribution in our study  
223 represents a single event location, all with equal weight, all component distributions are weighted equally, with the  
224 same number of samples drawn from each event location PDF in NonLinLoc ( $s_i = 5000$ , where  $s_i$  is the number of  
225 samples drawn for event index  $i = 1, \dots, n$ ).

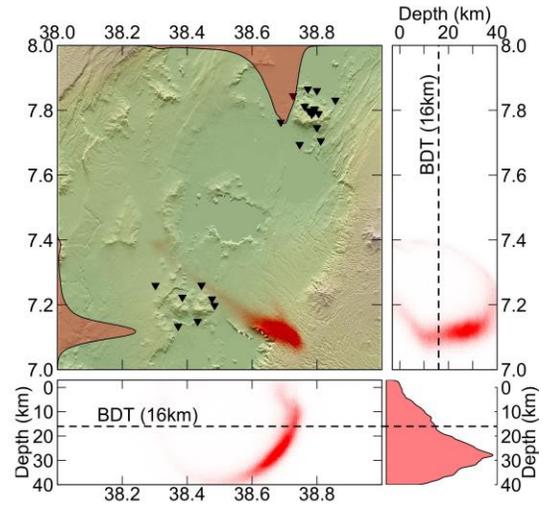
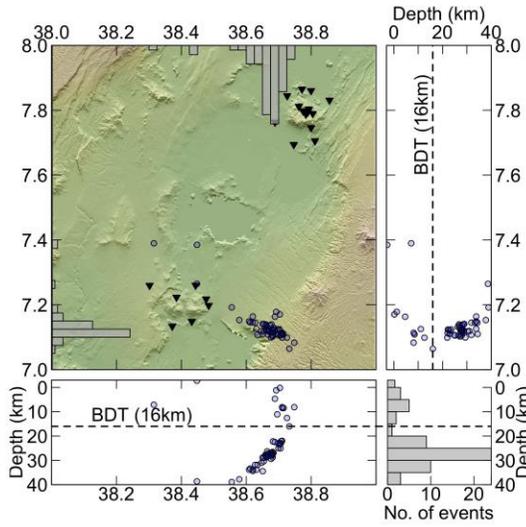
226  
227 Residuals between observed and expected travel times were greatly improved over initial manual picking  
228 through the CWT-WBS picking approach outlined in Section 4 (original manual-picking mean absolute travel-time  
229 residuals = 0.89 s; CWT-WBS mean absolute travel-time residuals = 0.23 s). However, reducing pick error did not  
230 markedly reduce the size of individual event location uncertainties and, in some cases, using a smaller number of  
231 higher confidence phase arrival times produced greater hypocentre location uncertainty (in terms of spatial spread)  
232 than using a larger number of lower confidence arrival times. This suggests that arrival time pick error has a smaller  
233 contribution to absolute location uncertainty in NonLinLoc than array geometry and estimates of travel time and  
234 velocity model error (Lomax et al., 2009). Despite the large uncertainty in individual event locations, the overall  
235 population (mixture) distribution for all event locations produces a clear, stable mode between 28 and 32 km depth  
236 for all velocity models and model error parameters (shaded density curves in Fig 7). Furthermore, the vast majority  
237 of event hypocenters are located between depths of 25 – 35 km (histograms in Fig 7) for any given level of Gaussian  
238 travel time error. This suggests that most events are indeed at lower crustal depths, although this becomes  
239 understandably less clear with very large travel time error levels (e.g.,  $\geq 10\%$  travel time error; Fig 7 bottom). The  
240 maximum likelihood estimates for hypocentre locations across all model runs also fall into two distinct clusters: a  
241 small, shallower cluster above the BDT zone (Muluneh et al., 2018) and a larger, deeper cluster between depths of  
242 20 – 35 km (Fig 7), with very few events located around the BDT zone itself.

243

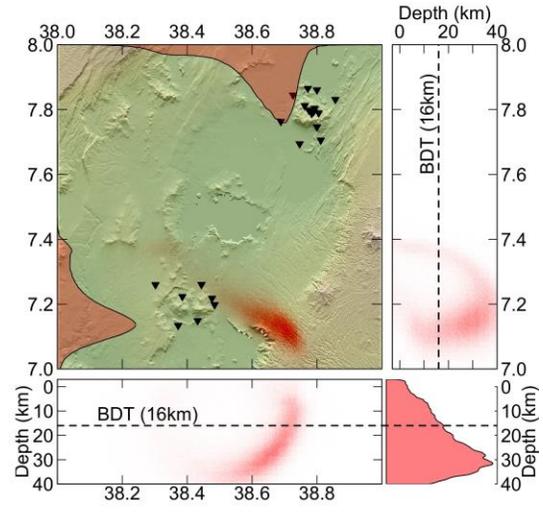
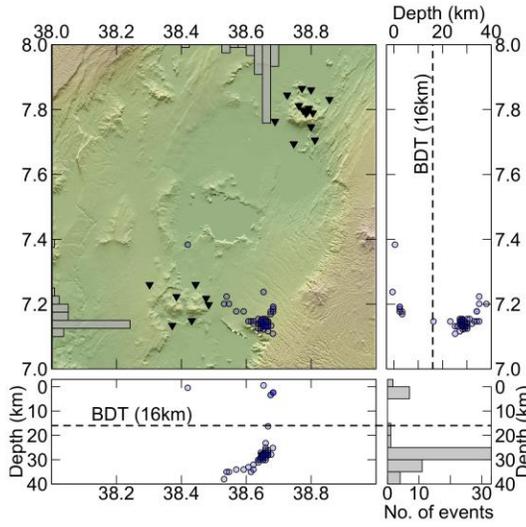
### Event Hypocenters

### Mixture Distributions

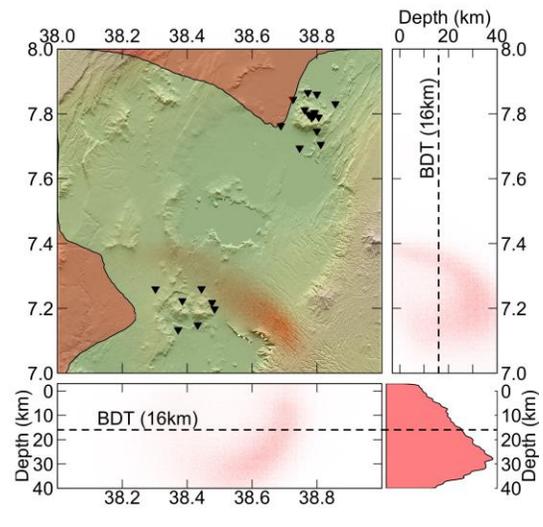
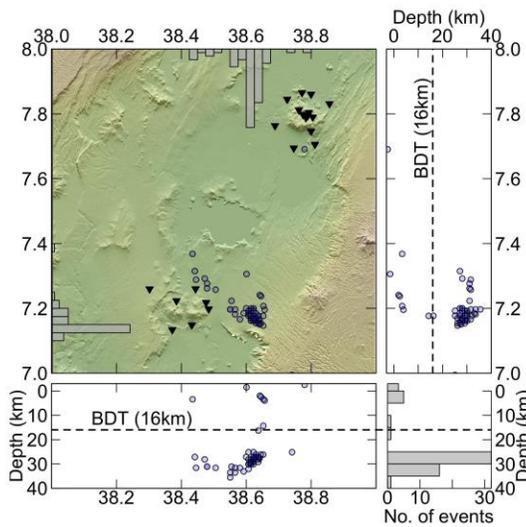
TT error:  
1 %



TT error:  
5 %



TT error:  
10 %



245 *Figure 7 – Event locations and mixture distributions for varying levels of travel time error (1% top, 5% middle, 10%*  
246 *bottom) using velocity Model 1 from Fig 5. All velocity models produced similar results. Left: Individual event*  
247 *locations using NonLinLoc (blue circles). Black triangles are seismic stations. Histograms show the number of*  
248 *events located within a 0.05 degree bin (latitude and longitude) or 5 km bin (depth). Right: Corresponding mixture*  
249 *distributions for all event locations (and estimates of their complete posterior PDFs). Red scatter and*  
250 *corresponding density curves represent mixture distributions of NonLinLoc event location PDFs with respect to a*  
251 *given plane (latitude, longitude, depth).*  
252

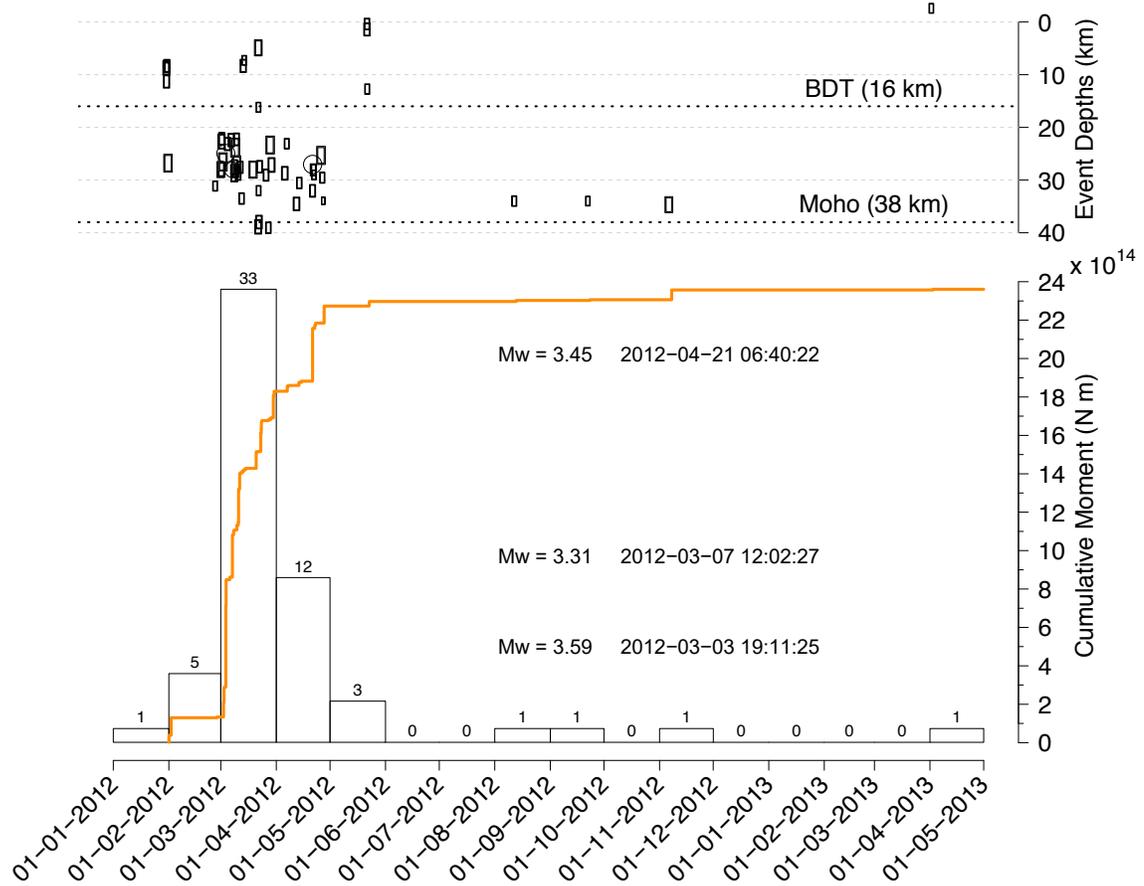
## 253 **6 Further Source Characteristics**

254

255 There is some evidence of repeating event source(s) from high inter-event cross-correlation (CC) values  
256 (e.g., Augliera et al., 1995) at station C02E, the closest station to the study area, with 47 out of a total of 58 P-wave  
257 arrivals having CC values  $> 0.7$  with at least one other event (4 sec window around P-wave arrival; at least 4 distinct  
258 multiplet groups identified). However, this is not seen at other stations across the two volcanic arrays, where inter-  
259 event CC values are consistently low due to low SNR, particularly across the Aluto array (on average,  $< 4$  out of 58  
260 events have CC value  $> 0.7$  with at least one other event at a given station). As such, it is difficult to determine  
261 whether events have similar source mechanisms or exploit any self-similarity in our analyses of event locations.

262

263 Most events occur over a two-month period (Fig 8), with 45 of 58 events and 90 % of cumulative seismic  
264 moment ( $M_0$ ) occurring during March and April 2012.  $M_0$  and moment magnitudes ( $M_W$ ) were calculated for each  
265 event at each station using spectral analysis of both P- and S-wave arrivals, with noise spectra subtracted, and Brune  
266 source model fitting (Abercrombie, 1995; Prejean & Ellsworth, 2001; Wilks, 2016). The values of  $M_0$  and  $M_W$  were  
267 then averaged across all stations to attain a final, single value of  $M_0$  and  $M_W$  for each event. To verify the quality of  
268 these predicted magnitude values, local magnitudes ( $M_L$ ) were also calculated, using a scale calibrated for the MER  
269 (Keir et al., 2006), with station corrections (determined by the average deviation of magnitudes measured at a given  
270 station) applied to account for the variability in the recording environments from station to station. Estimates of  $M_L$   
271 and  $M_W$  were found to be in close alignment, with the 2.5% and 97.5% quantiles for  $M_L - M_W$  equal to -0.28 and  
272 0.29, respectively. First motion polarities, where identified, for all events within the main cluster (January – June  
273 2012) show a consistent trend: first motions across one array were consistently opposite to those at the other. During  
274 this period, first motion was predominantly downward (dilation) across the Corbetti stations and upward  
275 (compression) across the Aluto stations, although these polarities reversed to their opposite sign on at least three  
276 occasions. However, the final four isolated events (between August 2012 and May 2013) had distinctly different  
277 behavior: the polarities were the same across both arrays, with two events having compression first motion across all  
278 stations and the other two events having dilation first motion across all stations. This overall behavior is interesting,  
279 particularly as these events are clustered in time, as it suggests they may represent different sources or a complex  
280 pattern from a single process.



281  
 282 *Figure 8 – Bottom: cumulative moment for all located events (orange line) and histogram of number of events per*  
 283 *month (yellow). 90% of total moment occurred in March/April 2012. Largest three events and moment size indicated*  
 284 *in blue. Top: Corresponding event depths (grey and blue circles) from Fig 7, TT (travel-time) error = 0.01. Moho*  
 285 *depth of 38 km approximated from Dugda et al. (2005) and Stuart et al. (2006).*  
 286

## 287 **7 Discussion**

288  
 289 Mixture distributions were used to overcome potential misinterpretations arising from large depth  
 290 uncertainties in individual hypocentre locations and to assess the overall likelihood of lower crustal events. The  
 291 overall mixture distribution mode and majority of maximum likelihood estimates (MLEs) for hypocentre locations  
 292 lie between 25 and 35 km depth for all velocity models and levels of travel-time error (Fig 7). These results indicate  
 293 that earthquakes beneath the magmatic MER and its border faults likely occur at lower crustal depths and far below  
 294 the generally recognized seismogenic zone along the rift (Keir et al., 2009; Muluneh et al., 2018; Yang & Chen,  
 295 2010).  
 296

297 Our MLE hypocentre locations suggest a possible bimodal distribution of event depths (above 15km and  
298 below 20 km), which is consistent with depth distributions previously observed near the MER (Keir et al., 2009) and  
299 in less magmatic sections of the EAR (Yang & Chen, 2010). The ephemeral and very localized nature of this  
300 seismicity (Figs 7 and 8) combined with low event magnitudes (range: 1.9 – 3.6  $M_w$ ; median: 2.5  $M_w$ ), the  
301 magmatic setting associated with the Corbetti volcano and MER, and the adjacent hot springs around the Wondo  
302 Genet scarp at the surface all suggest that these lower crustal events are likely related to fluid or magmatic processes  
303 (Keir et al., 2009; Yang & Chen, 2010) rather than slip on cold or modified crust.

304

305 Seismicity in the area around Wondo Genet, where our mixture location distribution mode lies, has not  
306 been observed in previous (e.g., Maguire et al., 2003) or subsequent (e.g., Lavayssière, Greenfield, et al., 2019)  
307 seismic deployments, with the latter study deploying a broadband seismometer directly above the area of seismicity  
308 identified in this paper. This lack of subsequent seismicity supports the interpretation that these events were, in fact,  
309 related to a single, ephemeral intrusion or transient exsolution / migration of volatiles, rather than ongoing volcanic  
310 or shallow hydrothermal activity associated with the Corbetti caldera or Wondo Genet hot springs. A reasonable  
311 interpretation from the identified pattern of first motion polarities across the two arrays (Section 6) could be a stable  
312 source mechanism during the main period of an intrusion (i.e., the main cluster of events during Jan – April 2012)  
313 followed by a more complex process following the intrusion event. Alternatively, the transient presence of hot fluids  
314 may have increased pore pressure or reduced the effective normal stress on the border and cross-rift faults at this  
315 intersection. Unfortunately, additional assessments of source mechanism (e.g., focal mechanism determination / full  
316 waveform inversion) and relative event locations (e.g., double-differencing and coda wave interferometry) all  
317 yielded poor or unstable solutions due to the low number of events, large array gap, low number of picks / stations,  
318 unknown velocity model error, low SNR and low cross-correlation values. As such, assessment of source must come  
319 from temporal and spatial characteristics combined with plausible physical mechanisms within the regional setting.

320

321 The weakness of our analysis lies in the relative location of instruments available during the study period,  
322 which yields large depth errors for individual hypocentre locations regardless of velocity model or error estimates  
323 (Lomax et al., 2009). As such, it is impossible to state whether any of these events lie below the Moho, assumed to  
324 be at 37 – 40 km depth for the southern MER (Ayele et al., 2004; Dugda et al., 2005; Stuart et al., 2006), or whether  
325 they are all constrained within the lower crust. Stable overall population (mixture) distributions, however, reveal a  
326 clear mode between 28 and 32 km depth, regardless of velocity model or parameter adjustments, and thus strongly  
327 suggest that at least some of these events are deep. Whilst the highly heterogeneous composition beneath the rift and  
328 these volcanic centres makes it difficult, or even impossible, to know which level of error is most appropriate, an  
329 error level of no more than 10 % of travel-time seems suitable given the error bounds published for one of the  
330 regional velocity models used in this study (approx. 1.6 % difference in total travel-time; Daly et al., 2008) and the  
331 maximum absolute difference in travel-times between all models used (approx. 8 %).

332

333 One way in which our hypocentre location estimates could be placed artificially deep is through use of a  
334 velocity model that is slower than the true Earth velocity structure (e.g., Poliannikov & Malcolm, 2016). Where the  
335 velocity model used is incorrect, the direct-search approach of NonLinLoc provides a better estimate of location  
336 hypocentre than linearised methods, as negative and positive travel time residuals need not be balanced to produce a  
337 complete posterior probability distribution. By contrast, linearised approaches produce a single point estimate, often  
338 with Gaussian errors subsequently calculated (Gesret et al., 2015). It is possible to use approaches which jointly  
339 infer the velocity structure and event origin parameters to obtain PDFs of the hypocentre with the uncertainty of the  
340 velocity included (e.g., Piana Agostinetti et al., 2015), but these require either good coverage of the volume being  
341 imaged or well-constrained prior information on velocities to tightly locate events, which is not available here.  
342 Regardless, every care has been taken to use velocity models representative of the region through use of several  
343 previously published models (Daly et al., 2008; Mackenzie et al., 2005) and a range of travel time error levels that  
344 are consistent with other published estimates of velocity structure (e.g., Keranen et al., 2009). Furthermore, the  
345 previously identified magmatic-hydrothermal activity beneath Aluto and Corbetti volcanoes (Lloyd, Biggs, Birhanu,  
346 et al., 2018; Wilks et al., 2017) would suggest that the true Earth velocity structure along the ray path to these  
347 volcanic centres would more likely be slower, rather than faster, than our models, which do not incorporate any  
348 adjustments for these volcanic centres.

349

350 The source of lower crustal seismicity and processes by which magmatism evolves within the crust in  
351 continental rift zones remains an outstanding question (e.g., Weinstein et al., 2017). The 2012 – 2013 seismicity east  
352 of Corbetti volcano appears to lie at a potential intersection between the rift border and a pre-existing cross-rift  
353 structure beneath the Corbetti caldera (Lloyd, Biggs, Wilks, et al., 2018). However, the apparent NW-SE  
354 linear/listric distribution of events away from Corbetti is almost certainly an artefact of array geometry (Lomax et  
355 al., 2009), with the vast majority of individual location PDFs from NonLinLoc marking out this cross-rift ‘trend’, so  
356 it is impossible to say whether this seismicity relates to or indicates the easternmost extent of this cross-rift structure.  
357 Event locations near Corbetti do, however, fit with observations of lower crustal seismicity in other regions of recent  
358 volcanism, both around the MER (Keir et al., 2009) and other volcanic centres (e.g., Neuberg et al., 2006; Soosalu et  
359 al., 2010). The distance of the overall distribution mode of seismicity at ~ 30 km from the Corbetti caldera suggests  
360 that the source of these events is exploiting a potential point of weakness along the rift border fault or cross-rift  
361 structure rather than being directly related to the magmatic storage processes beneath Corbetti (Lloyd, Biggs,  
362 Birhanu, et al., 2018; Lloyd, Biggs, Wilks, et al., 2018). Plausible mechanisms, given the temporal and spatial  
363 distribution of this seismicity, include the movement of magma and/or exsolution of volatiles causing transient high  
364 strain rates and pore fluid pressures that induce seismicity (e.g., Greenfield & White, 2015; Keir et al., 2009;  
365 Soosalu et al., 2010) or reduce the effective normal stress on the border or cross-rift faults (e.g., Reyners et al.,  
366 2007), as opposed to an unusually strong lower crust (e.g., Craig et al., 2011). Further seismic monitoring, both in  
367 this area and of Ethiopian volcanoes in general, would provide greater opportunity to observe such lower crustal  
368 events again in the future, constrain source process and identify how magma migrates from mantle to crust in  
369 continental rift zones.

370

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372

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 380 on data access and availability. Velocity models, phase arrival pick times, a catalog of events and an example link to  
 381 download SAC waveform data through IRIS can be found in the Supplementary Material.

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387

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